Equatorial waves and superrotation in the stratosphere of a Titan general circulation model

Neil Lewis¹, Peter Read¹, Nick Lombardo², Juan Lora²

¹Atmospheric, Oceanic and Planetary Physics, University of Oxford ²Department of Earth and Planetary Sciences, Yale University

FDEPS Seminar: 1 December 2023

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

1 Key features of Titan

Parameter	Titan	Earth
★ Radius (km)	2575	6371
\star Rotation period (Earth days)	16	1
\star Surface gravity (m s-2)	1.35	9.81
\star Surface pressure (bar)	1.5	1
Atmospheric composition	94% N ₂ 5% CH ₄	78% N ₂ 21% O ₂ ~.1% H ₂ O
\star Insolation (W m-2)	14.9	1368
★ Obliquity (deg)	27	23.45
Eccentricity	0.029	0.0167
\star Length of year (Earth years)	29	1





Key features of Titan

- Complicated radiative budget and T profile
 - Greenhouse-like troposphere
 - Heated via surface
 - Anti-greenhouse stratosphere
 - Heated directly via hazes
- Complicated chemistry and thermodynamics
 - Optically thick photochemical hazes in stratosphere
 - Condensible constituents (CH₄) with clouds, rain, lakes....
- Strong seasonal variability





- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

2 Equatorial superrotation – what is it?

Super-rotation is a phenomenon in atmospheric dynamics where the axial angular momentum of an atmosphere in-some-way exceeds that of the underlying planet.

A maximum off of the equator will typically be inertially unstable, so superrotation is invariably manifest as a zonal jet centred on the equator.

Hide's theorem says that in axisymmetric, inviscid atmosphere

$$\frac{\mathrm{D}\overline{m}}{\mathrm{D}t} = 0$$

where $m = a \cos \vartheta (\Omega a \cos \vartheta + u)$, and the overline denotes a zonal average.

Friction + overturning will lead to m = 0. Therefore, some process that can transport m up-gradient is required. This can be achieved by eddies.

2 Equatorial superrotation – results from idealised modeling

Idealised 'Earth-like' models show that superrotation emerges naturally for planets with low rotation rate / radius.



Smaller / slower rotating

2 Equatorial superrotation – results from idealised modeling

Idealised 'Earth-like' models show that superrotation emerges naturally for planets with low rotation rate / radius.

Barotropic, ageostrophic instability generates a 'Rossby—Kelvin wave' that induces equatorward momentum flux *during spin-up*:



Superrotation maintained by 'equatorial Rossby waves', generated by baroclinic/barotropic instability:



 Situation on Titan is complicated by seasonal cycle and latitudinally asymmetric shear



- Situation on Titan is complicated by seasonal cycle and latitudinally asymmetric shear
- Leads to a seasonal cycle in the momentum budget



- Situation on Titan is complicated by seasonal cycle and latitudinally asymmetric shear
- Leads to a seasonal cycle in the momentum budget
- What are the eddies like?
- Which contribute most to AM transport?
- where do they come from?
 - Regular barotropic instability?



- Situation on Titan is complicated by seasonal cycle and latitudinally asymmetric shear
- Leads to a seasonal cycle in the momentum budget
- What are the eddies like?
- Which contribute most to AM transport?
- where do they come from?
 - Regular barotropic instability?



From Lebonnois et al. [2012]

- Situation on Titan is complicated by seasonal cycle and latitudinally asymmetric shear
- Leads to a seasonal cycle in the momentum budget
- What are the eddies like?
- Which contribute most to AM transport?
- where do they come from?
 - Regular barotropic instability?



(c) $L_s = 295^\circ$

From Lebonnois et al. [2012]

- Situation on Titan is complicated by seasonal cycle and latitudinally asymmetric shear
- Leads to a seasonal cycle in the momentum budget
- What are the eddies like?
- Which contribute most to AM transport?
- where do they come from?
 - Regular barotropic instability?

20 -atitude (deg N) -20 30 60 90 120 150 210 240 270 300 Planetocentric solar longitude (deg Ls) -1.44e-11 -8.00e-12 -1.60e-12

-2.08-11 -1.44e-11 -0.00e-12 -1.00e-12 4.00e-12 -2.40e-11 -1.76e-11 -1.12e-11 -4.80e-12 1.60e-12 8.00e-12

330

Contours of $\frac{\partial^2 \overline{u}}{\partial v^2} - \beta$ & plot of $\partial M / \partial t$

From Newman et al. [2011]

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

3 Open questions + aim

Which (equatorial) wave modes are supported in Titan's stratosphere, and what is their spatial structure?

Which contribute to maintaining equatorial superrotation?

What generates them?

What relevance do the results from idealised modelling have?

- Theoretical understanding
- Observed behaviour of Titan's climate?

Our aim: to answer some of these!

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

4 Model configuration

Analyse equatorial waves and (Lombo eddy angular momentum fluxes in TAM

TAM = Titan Atmospheric Model

GFDL spectral core with T21 horizontal resolution (~4.5 degree)

50 levels in vertical 0 to ~450km

Realistic RT with seasonal cycle and absorption of solar radiation due to haze

Methane hydrological cycle

Data coverage from L_{sol} 180° – 360°, output is 6hrly average (NH autumn equinox -> NH winter -> NH spring equinox)

TAM Model configuration

 10^{-2}

 10^{-1}

10⁰

10¹

10²

10³

60

Pressure (mbar)

- GFDL spectral dynamical core
 - Cf ISCA model
- T21 horizontal resolution
 - ~4.5° resolution in lat x lon
- 50 vertical levels
 - 0 < z < 450 km
- Realistic radiative transfer
 - seasonal cycle,
 - solar absorption due to haze
 - IR cooling due to hydrocarbons and CIA due to $N_{\rm 2}$ etc.
- Mellor-Yamada boundary layer
- Betts-Miller convection
- CH₄ hydrological cycle
- Variable chemical composition [potentially]

Lombardo & Lora [2023]

- L50 - L32

-HASI

180

120

Temperature (K)

90

150

NH early Winter

NH Spring

NH early Summer

midsummer

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

5 Basic circulation and eddy momentum fluxes

Strongly stratified (N>0), plus N depends strongly on altitude / pressure:

Seasonal cycle of u, Psi (TEM), and EP fluxes:

5 Basic circulation and eddy momentum fluxes

Key features:

Seasonal cycle of u, Psi (TEM), and EP fluxes:

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

Spectral analysis of equatorial waves – late autumn 6

propagation

(Andrews et al. 1987, Imamura 2006)

Both: waves propagate upwards

6 ... analysis of equatorial waves – late autumn,

upper stratosphere

Spectral decomposition of eddy momentum flux and geopotential height

6 ... analysis of equatorial waves – late autumn,

upper stratosphere

Horizontal structure for selected waves

6. ... analysis of equatorial waves – late autumn, lower stratosphere

Acceleration in lower stratosphere due to retrograde waves

6 ... analysis of equatorial waves – late autumn, lower stratosphere

Spectral decomposition of eddy momentum flux and geopotential height

Phase speed [m s⁻¹]

a) $\overline{z'^2} \times p/p_r$ $L_{\rm s} = 230^{\circ}, p = 16.1 \, \rm hPa$ Prograde and retrograde waves present 512 at equator, again! 128 25Latitude 32 Ē Acceleration at equator due to modes with low intrinsic frequency, again! 100 150 200 Phase speed [m s⁻¹] - BUT, mostly retrograde b) $L_{\rm s} = 230^\circ, p = 16.1 \, {\rm hPa} - \rho_0 \frac{\partial \overline{m'v'\cos\vartheta}}{\partial \cos\vartheta\partial\vartheta}$ 0.56 Note also signature of eastward waves, and westward $10^{-3} \text{ m}^2 \text{ s}^-$ 0.24 25 wave with cp=-12 ms-1, similar to upper stratosphere Latitude -0.24 × -25-50What do they look like? -0.56 150100 200

6 ... analysis of equatorial waves – late autumn,

6 ... analysis of equatorial waves – late autumn, lower stratosphere

What about vertical structure?

Spectral decomposition of eddy geopotential (at equator):

A number of waves have coherent vertical structure.

Evidence for a critical region through which disturbances can not propagate vertically

Prograde waves in lower stratosphere appear to be concentrated onto certain pressure levels.

Roughly coincident with sharp gradient in background wind and also buoyancy freq.

6 ... analysis of equatorial waves – late autumn,

lower stratosphere

What about vertical structure?

What do they look like?

6 ... analysis of equatorial waves – late autumn,

lower stratosphere

What do they look like?

5 ... analysis of equatorial waves – generation mechanisms

Ageostrophic, shear instability?

Occurs when
$$\operatorname{Fr} \equiv \frac{\omega_{\operatorname{Rossby}}}{\omega_{\operatorname{Kelvin}}} \approx \frac{\overline{u}_{\operatorname{ml}}/\cos\vartheta_{\operatorname{ml}}}{\overline{u}_{\operatorname{eq}} + \omega_{\operatorname{Kelvin}}^{+}/k} \approx 1 - 2$$

$$\omega_{\operatorname{Kelvin}}^{+} = -\frac{Nk}{[m^{2} + 1/(4H^{2})]^{1/2}},$$

5 ... analysis of equatorial waves – generation mechanisms

16.0 hPa

What about for retrograde waves, that exist all year round?

Barotropic instability? (cf. Lebonnois et al., Newman et al.) $\partial \overline{Q} / \partial \vartheta = 0$ 50 Latitude 0 -50180 225 320 270 360 Lsol 0.44 hPa 3.0 hPa 2.0 hPa

6.0 hPa

.0 hPa

Criterion met all year round

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

- 1 Key features of Titan
- 2 Equatorial superrotation
- 3 Open questions + aim
- 4 Model configuration
- 5 Basic circulation and eddy momentum fluxes
- 6 Spectral analysis of equatorial waves
- 7 Summary / discussion

7 Summary / discussion

Which (equatorial) wave modes are supported in Titan's stratosphere, and what is their spatial structure? - many: K, RK, ER, MRG (IG waves not shown)

Which contribute to maintaining equatorial superrotation?

- mostly RK in upper strat, ER in lower strat

What generates them?

- consistent with ageostrophic shear instability, and barotropic instability

What relevance do the results from idealised modelling have?

- some!

Despite differences, Titan resembles idealized model spin-up in upper stratosphere, and equilibrated idealized model in lower stratosphere....

Qs: Are waves properly resolved in vertical?

What wave modes do we expect (e.g., from instability analysis) from an asymmetric background with strong vertical shear?

Is the vertically trapped structure real?

Should there be more vertical momentum transport? (cf. the wind drop-off region on real Titan)

Situation on Titan complicated by seasonal cycle and latitudinally asymmetric shear

This leads to a seasonal cycle in momentum budget...

What are the eddies like?

(f) Ls=331.2

101

10°

10

10⁴

Pressure (Pa)

m/s

500

Altitude (km

120

100

These authors suggest regular barotropic instability as origin for the waves