

Torsional Alfvén waves in a dipolar magnetic field: *experiments and simulations*

Z. Tigrine ^{1,2}, **H-C. Nataf** ¹, N. Schaeffer ¹
P. Cardin ¹ and F. Plunian ¹

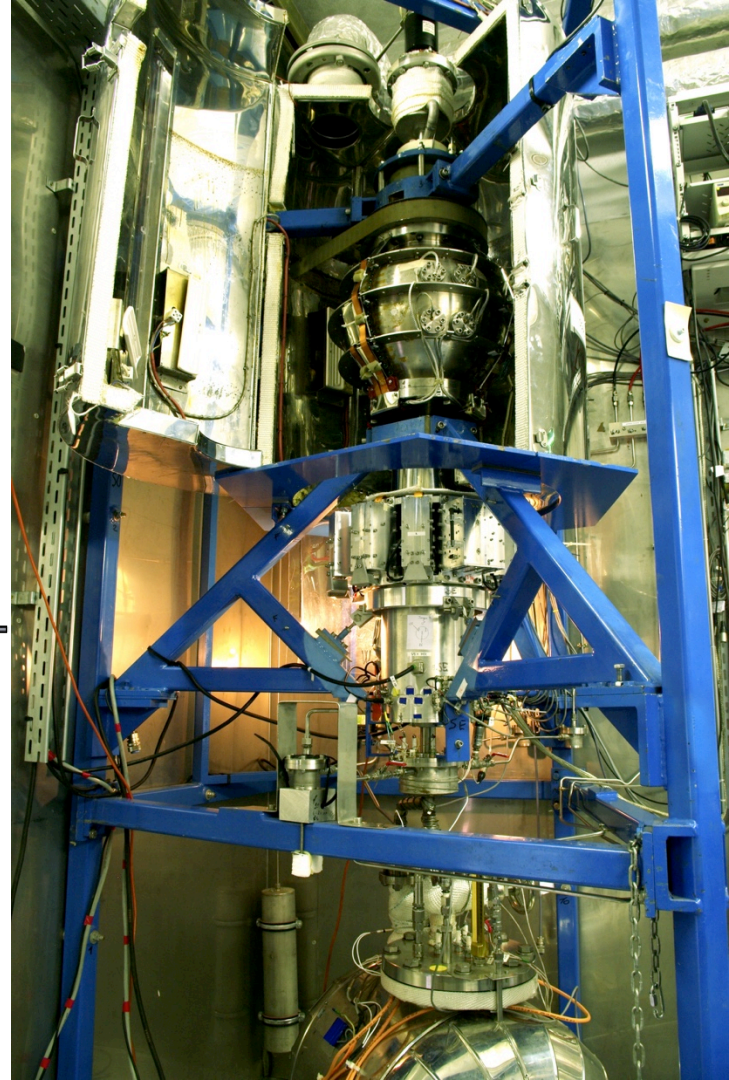
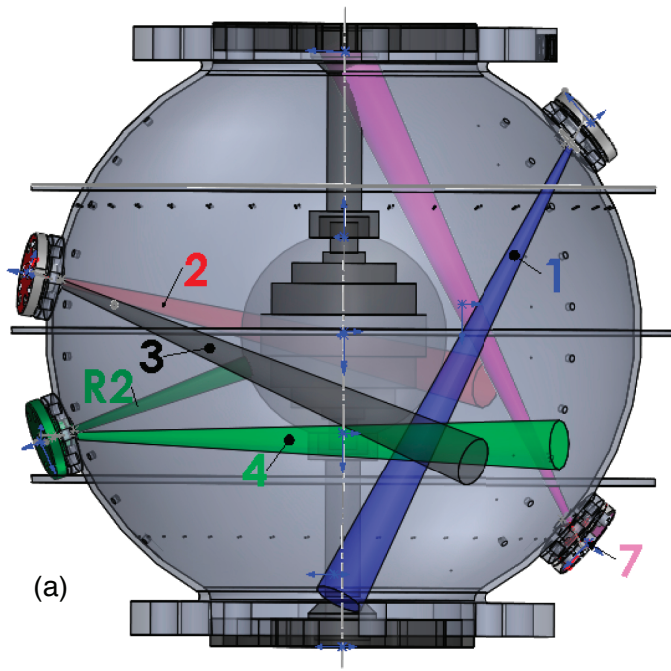
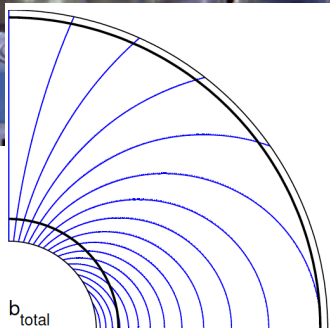
¹ *Univ Grenoble Alpes, CNRS, ISTerre, F-38000 Grenoble, France*

² *Univ des Sciences et de la Technologie Houari Boumediene USTHB, Algiers, Algeria*

The liquid sodium DTS experiment

- The goals of DTS:
 - explore the *magnetostrophic* regime, in which the Coriolis force and the Lorentz force are comparable.
 - Pave the way for a spherical Couette dynamo experiment.
- First measurements in 2005.
- DTS- Ω version in 2015.

The DTS set-up



Physical properties of liquid sodium

Property	symbol	Value	Unit
Density	ρ	930	kg.m ⁻³
Electrical conductivity	σ	9.10 ⁶	$\Omega^{-1} \text{ m}^{-1}$
Kinematic viscosity	ν	6.5. 10 ⁻⁷	m ² /s
Magnetic Diffusivity	η	8.84.10 ⁻²	m ² /s
Temperature	T	130	°C

Physical parameters of DTS

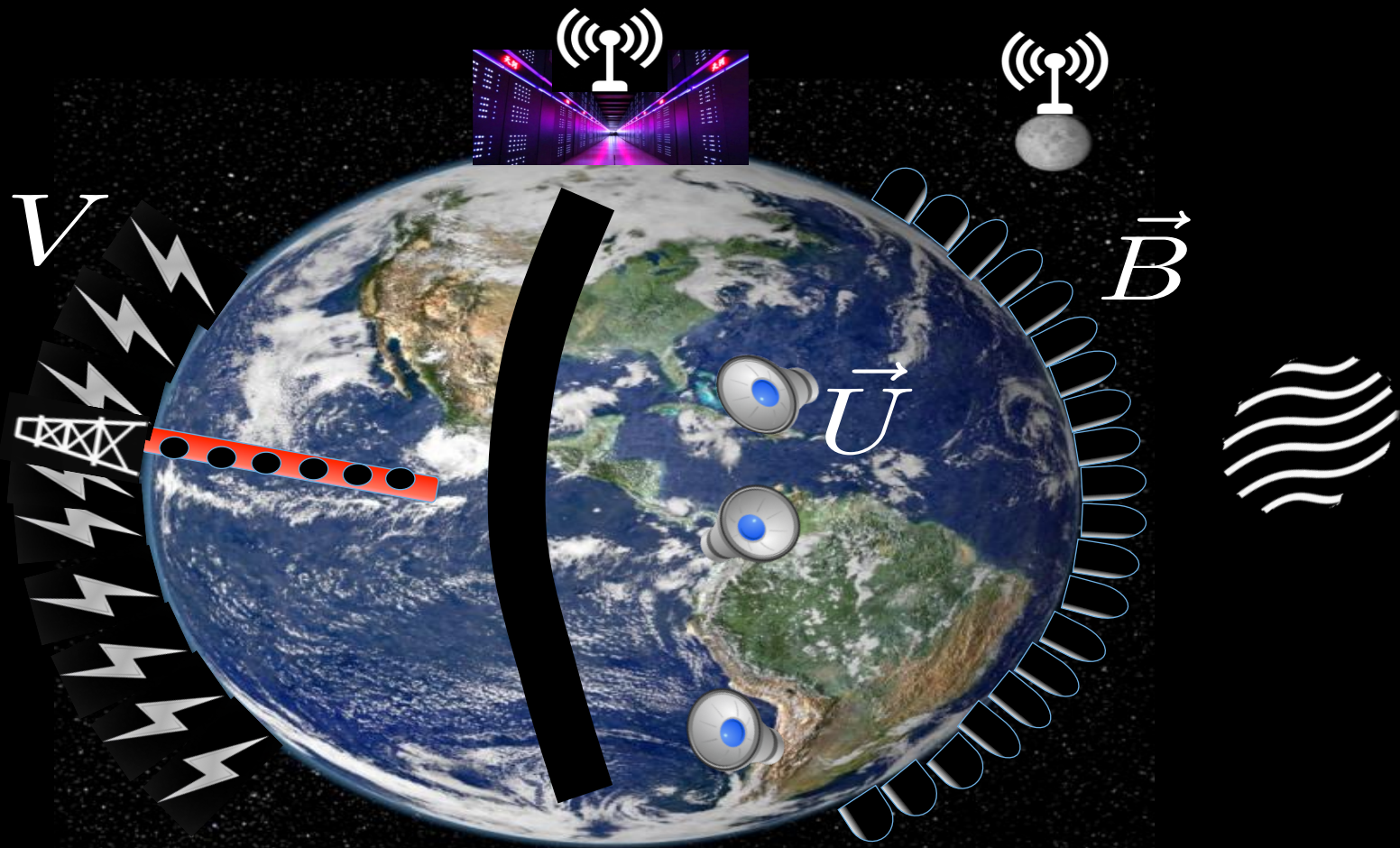
Property	symbol	Value	Unit
Outer sphere radius	r_o	210	mm
Inner sphere radius	r_i	74	mm
Outer sphere maximum rotation rate	f_o	15	Hz
Inner sphere maximum rotation rate	f_i	30	Hz
$B(r=r_o, \theta=\pi/2)$	B_o	8	mT
$B(r=r_i, \theta=\pi/2)$	B_i	173	mT

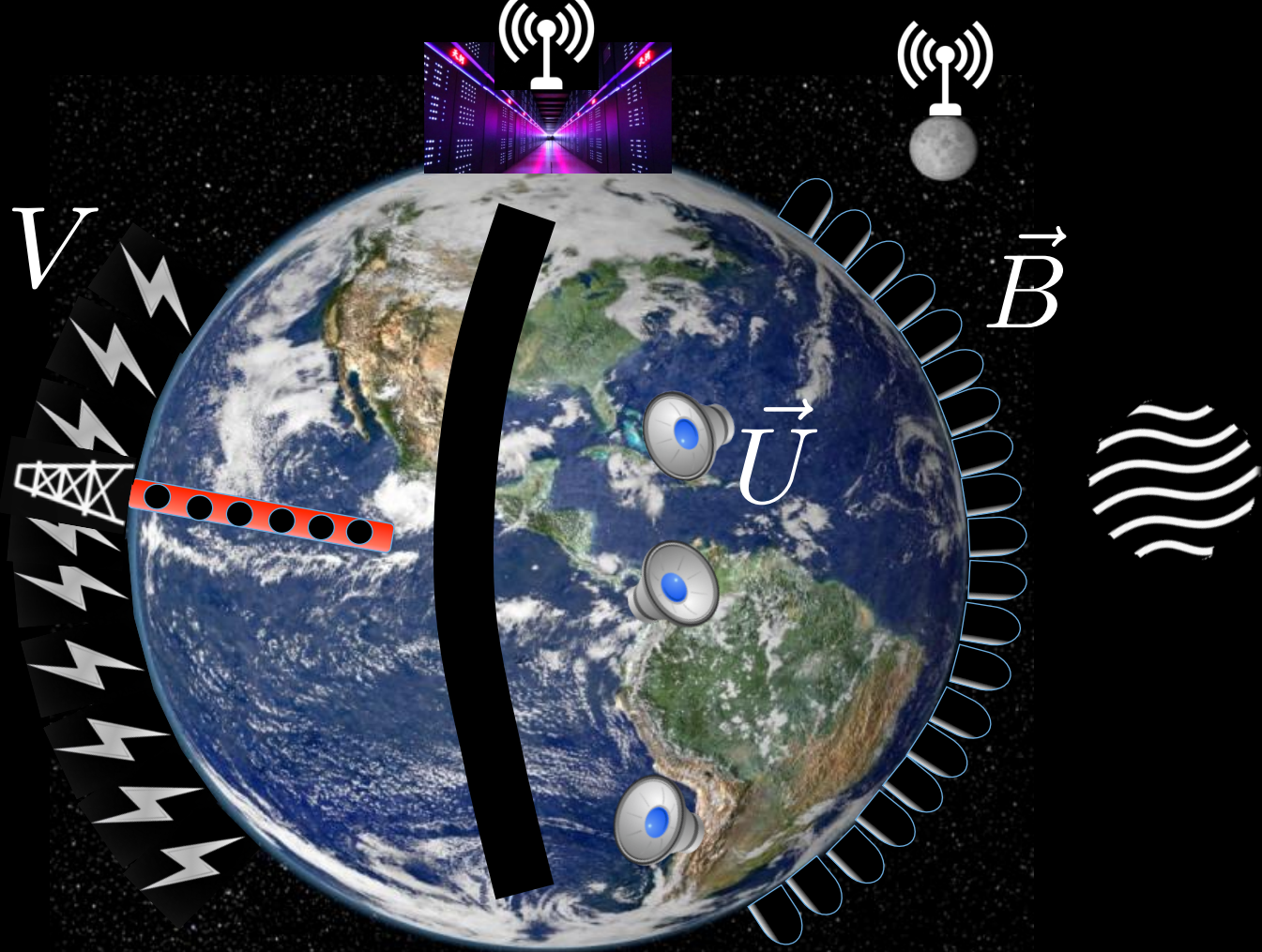
Dimensionless numbers for $f_o = 15$ Hz, $\Delta f = 30$ Hz

Number	expression	value	<i>Earth core</i>
Magnetic Prandtl	ν / η	7.4×10^{-6}	$\sim 10^{-5}$
Ekman	$\nu / \Omega r_o^2$	1.6×10^{-7}	$\sim 10^{-15}$
Reynolds	$\Delta \Omega r_o^2 / \nu$	1.3×10^7	
Magnetic Reynolds	$\Delta \Omega r_o^2 / \eta$	94	$\sim 10^3$

Measurement techniques

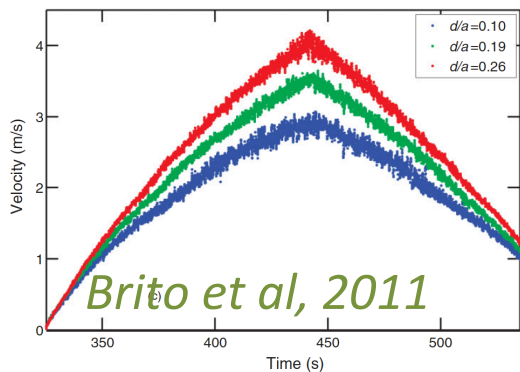
- Magnetic field at the surface
- Magnetic field in a sleeve inside the fluid
- Electric potentials at the surface
- Ultrasound Doppler Velocimetry



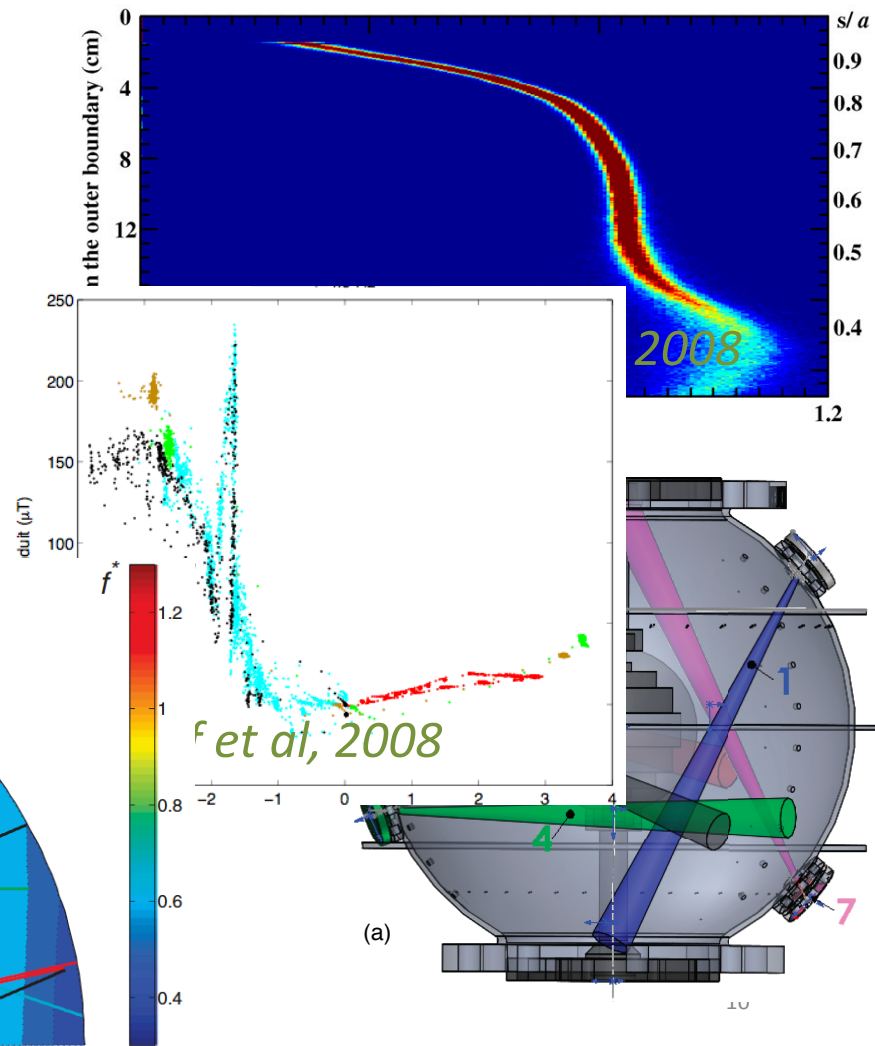
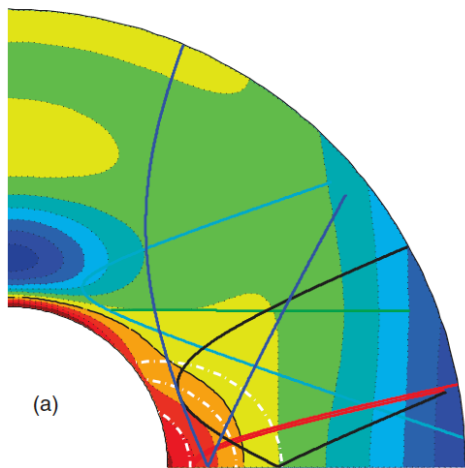


Main results

- Mean state:
 - Modified Taylor state
 - Induction peak at $Ro_{\text{eff}} \sim -1$
 - Super-rotation
 - $\Lambda \sim 1$ frontier

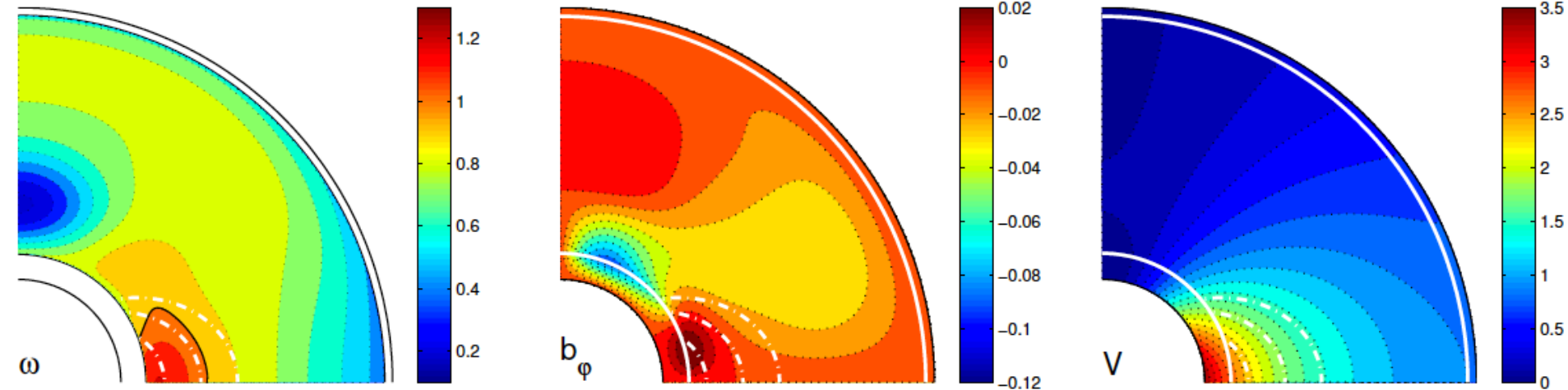


November 30, 2018



Main results

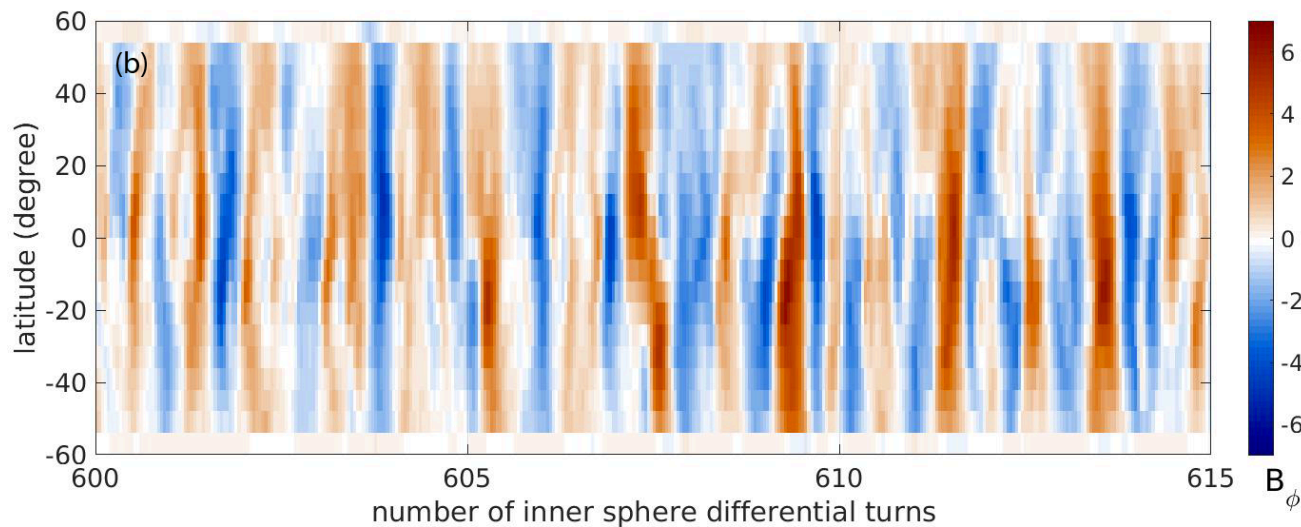
- Mean state:
 - Induction and diffusion: using the DTS experiment as a Navier-Stokes solver!



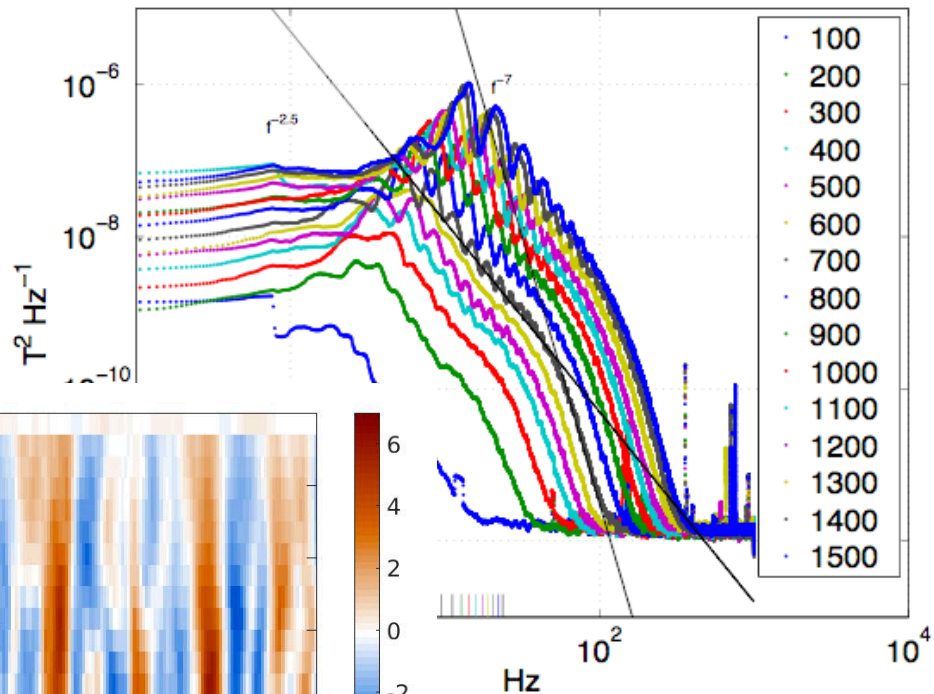
Nataf, 2013

Main results

- Fluctuations:
 - Modes and filaments



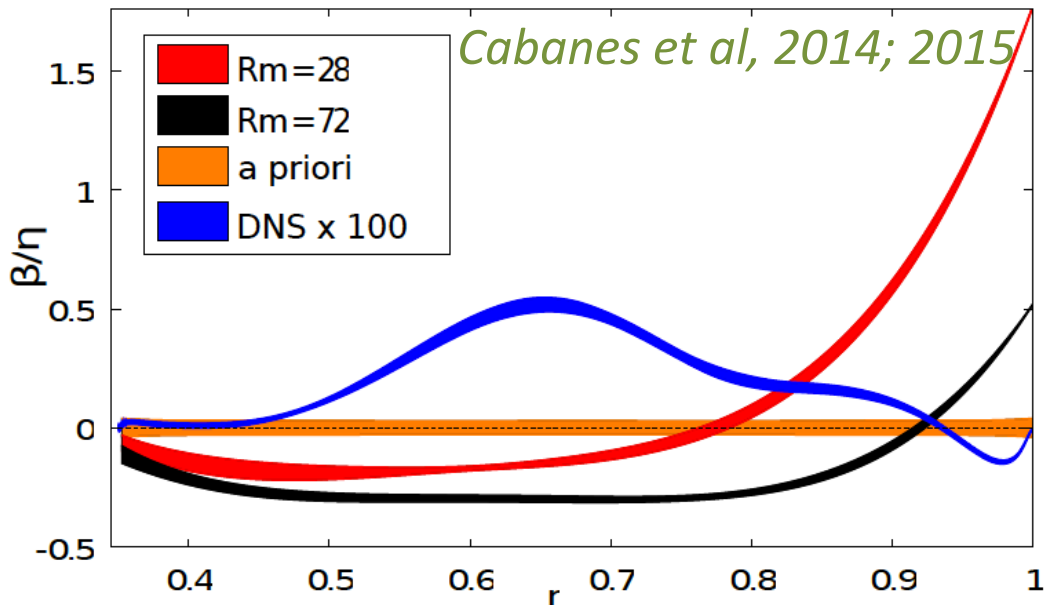
Magnetic energy spectrum, run 2 and 3, 31-01-06



Schmitt et al, 2008; 2013

Main results

- Fluctuations, diffusion, and mean flow:
 - Turbulence *reduces* magnetic diffusivity



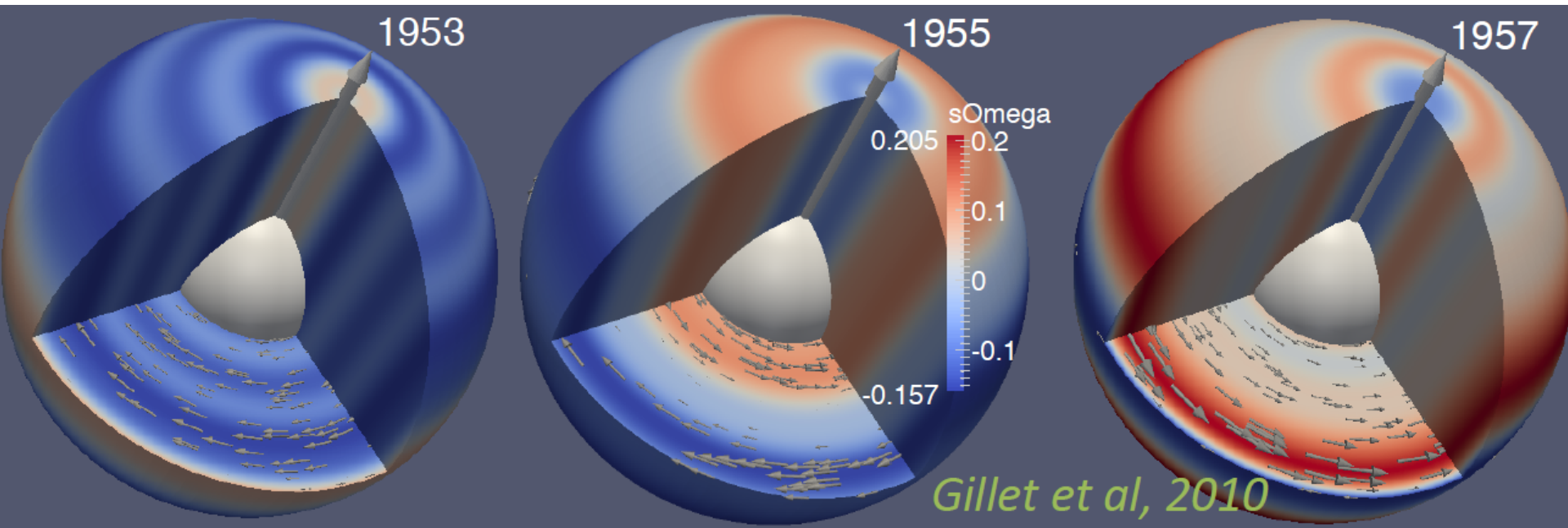
Torsional Alfvén waves in a dipolar magnetic field: *experiments and simulations*

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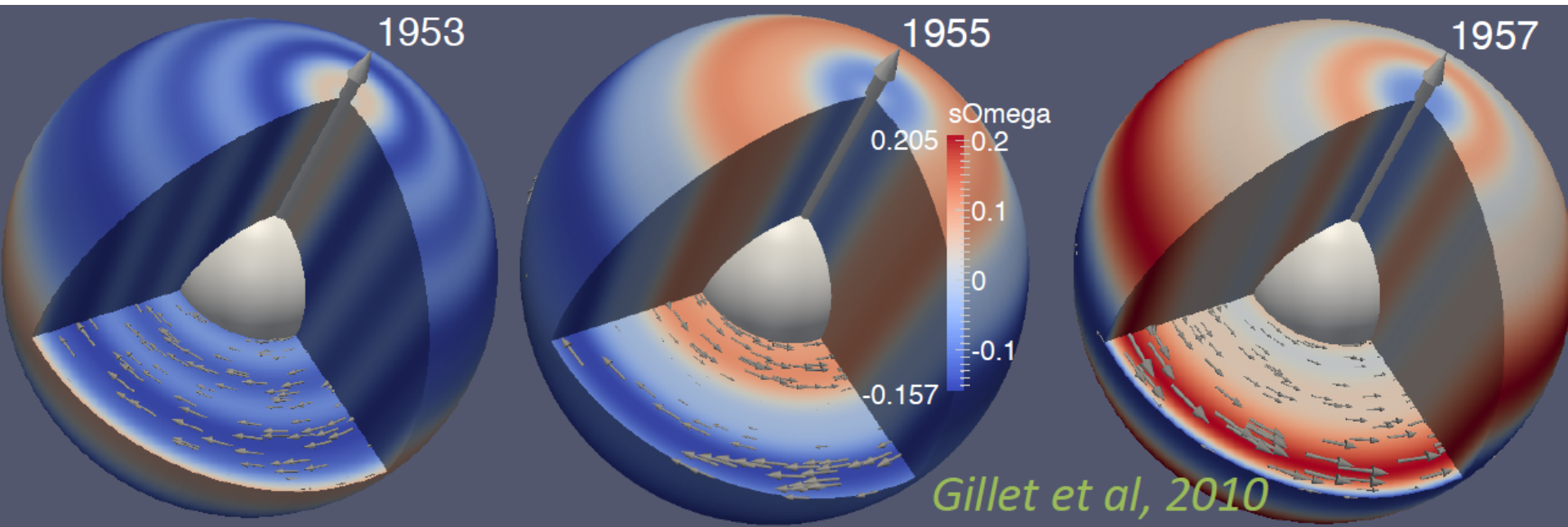
² *Univ des Sciences et de la Technologie Houari Boumediene USTHB, Algiers, Algeria*

Torsional Alfvén waves in the Earth's core



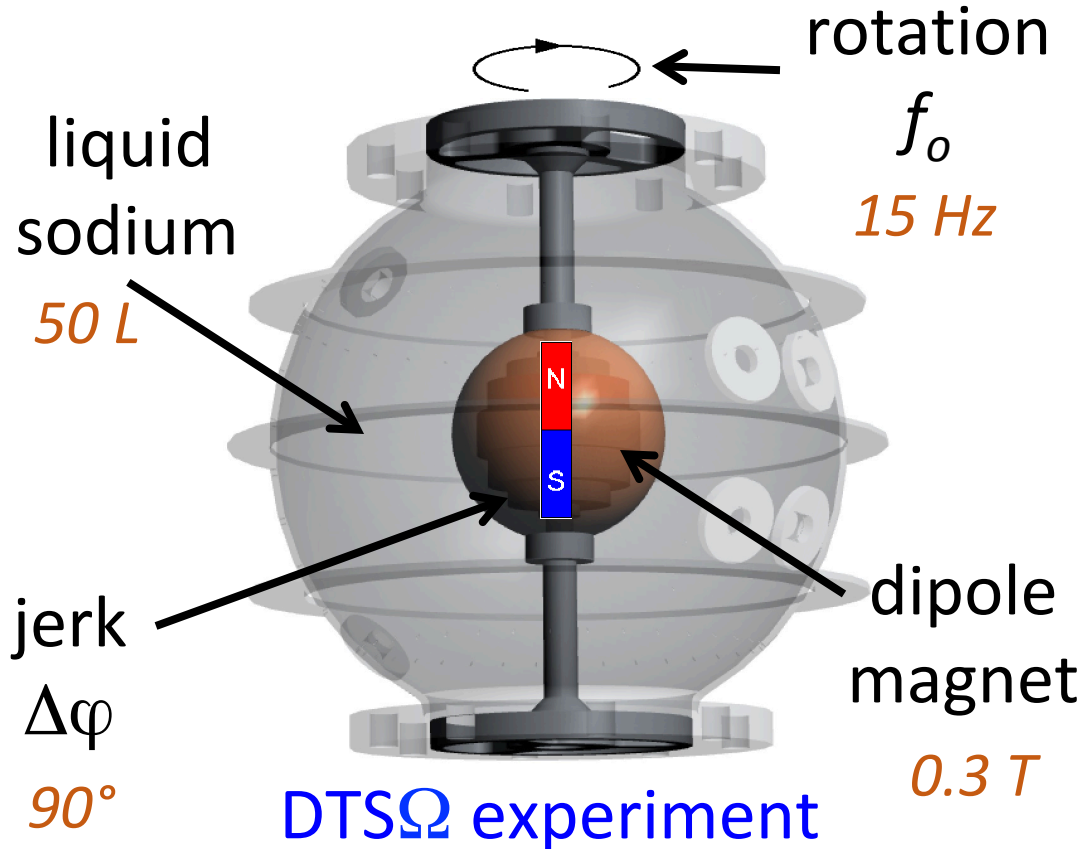
- Alfvén waves are strongly modified by the Coriolis force in planetary cores:
 - Alfvén waves that violate the Proudman-Taylor constraint are inhibited.
 - **Geostrophic Alfvén waves**, which are called torsional Alfvén waves, are favoured.

Torsional Alfvén waves in the Earth's core



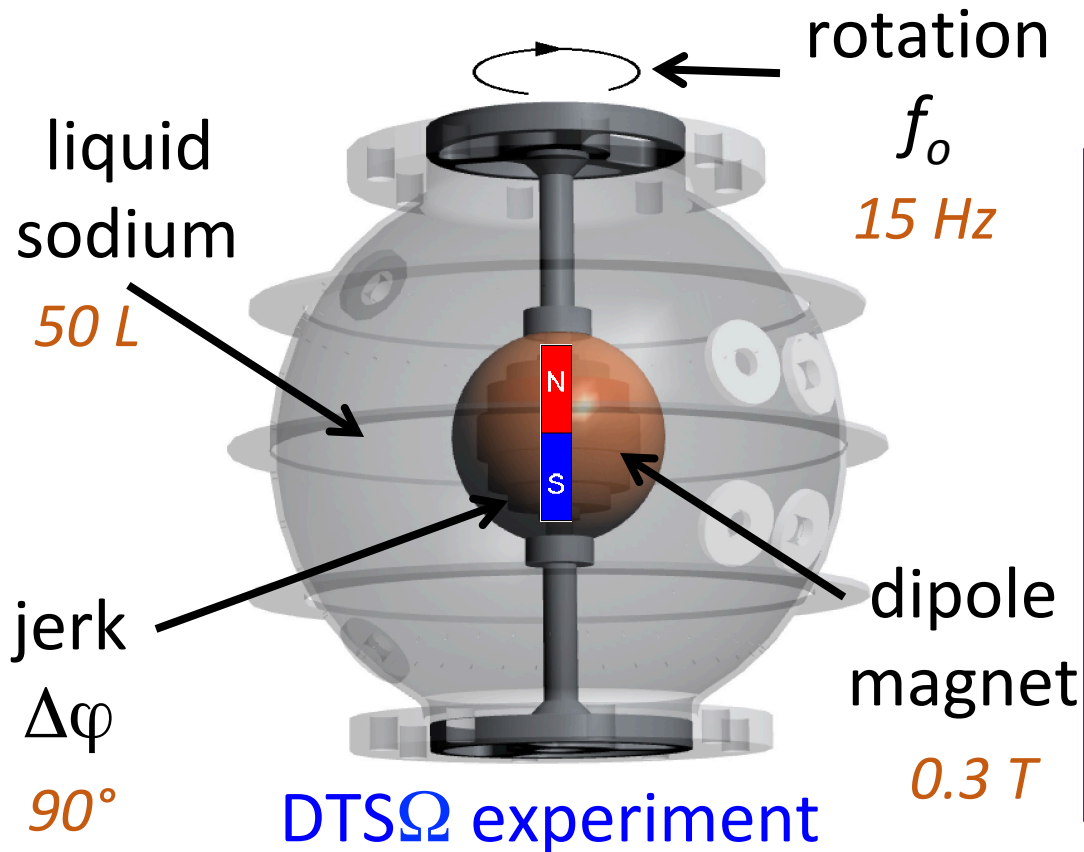
Alfvén time: $\tau_A = \frac{r_o}{V_A} = \frac{r_o \sqrt{\mu_0 \rho}}{B} \sim 4 \text{ years}$

Torsional Alfvén waves in the Laboratory...



Nataf et al, 2008
Brito et al, 2011
Cabanes et al, 2014

Torsional Alfvén waves in the Laboratory...



...and in the computer

XSHELLS software

Schaeffer, 2013

Figueroa et al, 2013

MHD axisymmetric

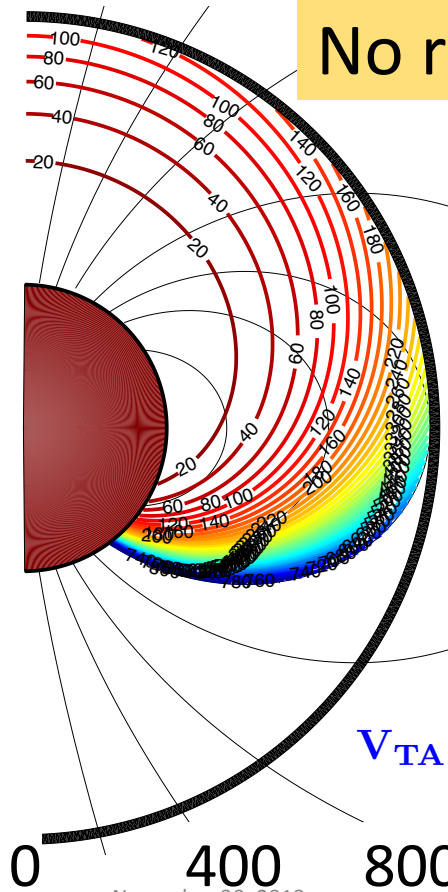
550 radial points, $I_{\max}=120$

adaptive time-step

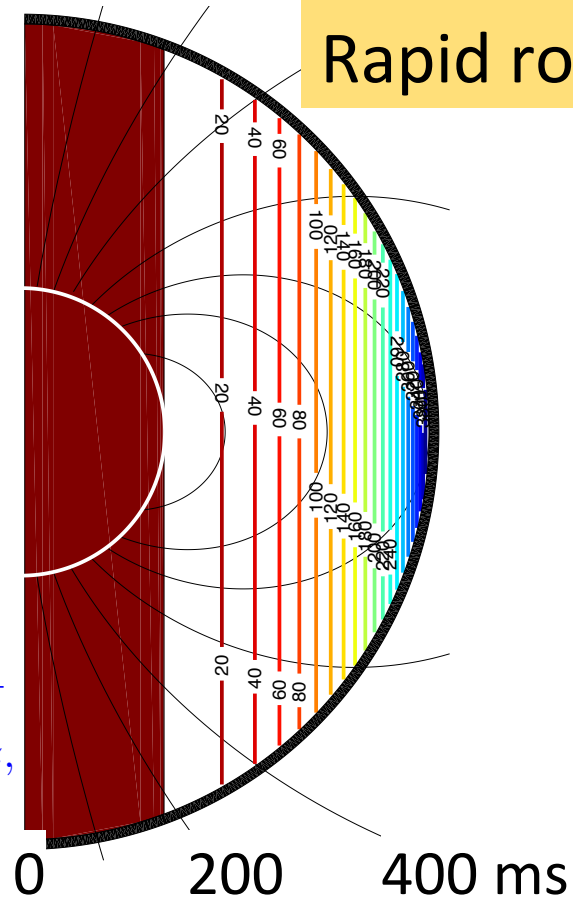
DTSΩ-like jerks

Wave fronts of *ideal* Alfvén waves in DTSΩ

No rotation



Rapid rotation



$$V_{TA}(s) = \hat{s} \sqrt{\frac{1}{2h(s)\mu_0\rho} \int_{-h}^h B_s^2(s, z) dz},$$

0 400 800 ms

November 30, 2018

FDEPS Research Seminar, Kyoto

0 200 400 ms

Wave fronts of *ideal* Alfvén waves in DTSΩ

No rotation

Lehnert number

$$Le = \frac{\tau_{\Omega}}{\tau_A}$$

$$\tau_{\Omega} = \Omega^{-1}$$

$$\tau_A = \frac{r_o \sqrt{\mu_0 \rho}}{B}$$

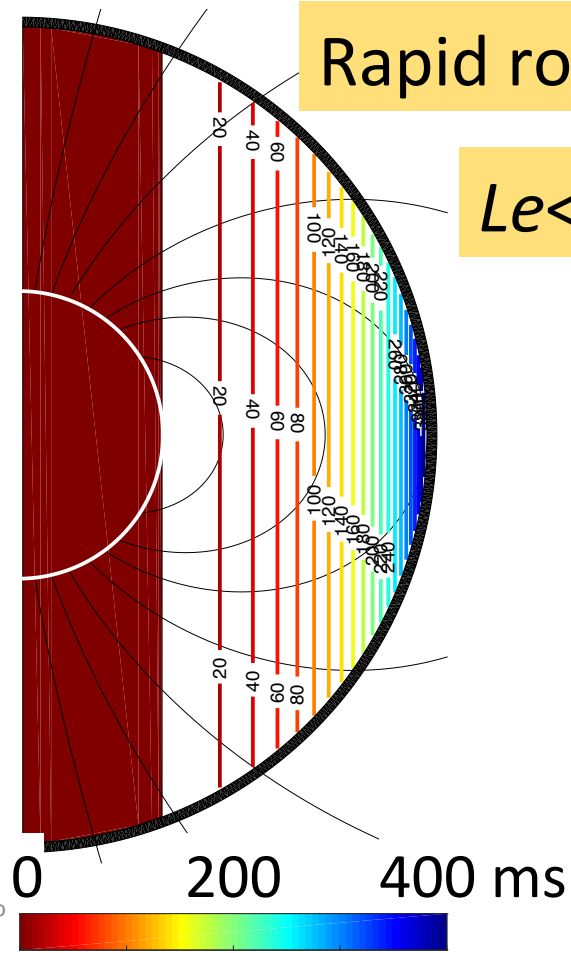
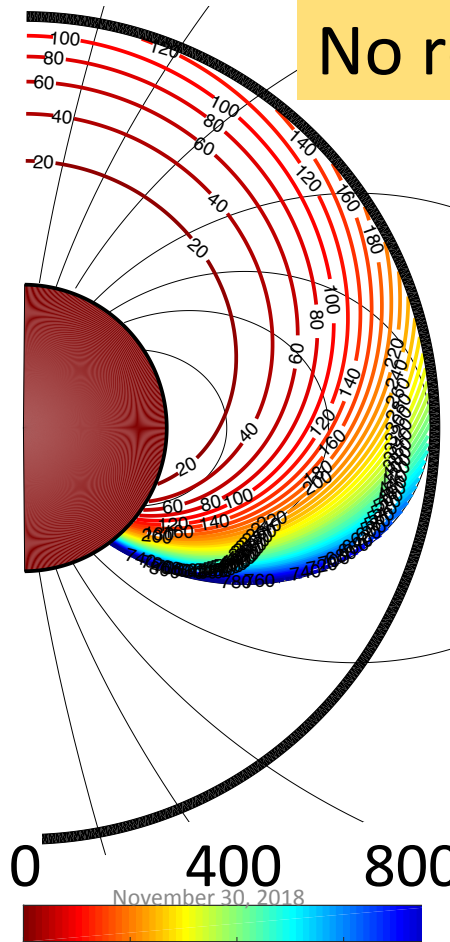
Lehnert, 1955

Jault, 2008

FDEPS Research Seminar, Kyoto

Rapid rotation

$Le \ll 1$



Magnetic diffusion

- Alfvén waves are very difficult to study in the lab because of the **large magnetic diffusivity** of liquid metals.

Magnetic diffusion time:

$$\tau_{\eta} = \frac{r_o^2}{\eta} = 500 \text{ ms}$$

Lundquist number

$$Lu = \frac{\tau_{\eta}}{\tau_A}$$

Lundquist, 1949

Lehnert, 1953

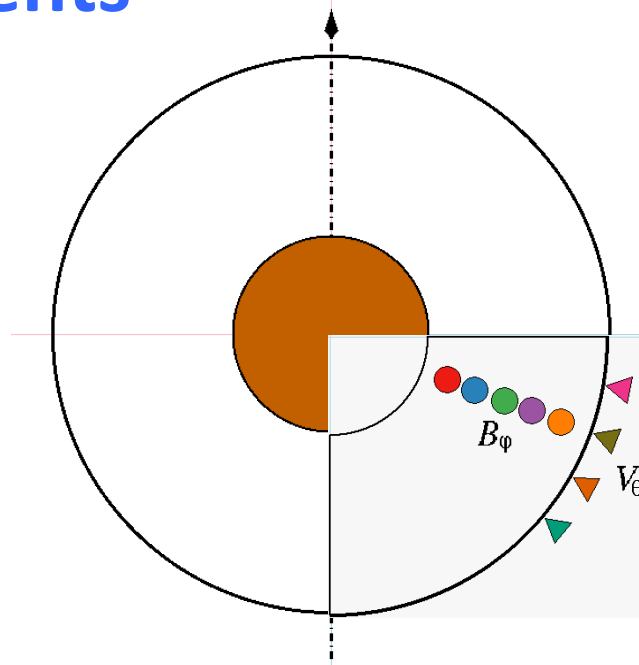
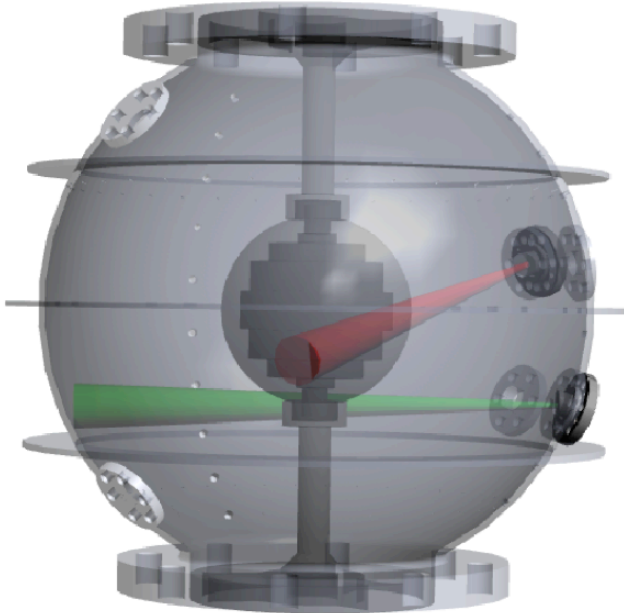
Jameson, 1961; 1964

Alboussière et al, 2011

Dimensionless numbers for $f_o = 15$ Hz

Number	expression	Inner sphere	Outer sphere	<i>Earth core</i>
Lehnert	τ_{Ω} / τ_A	0.25	0.01	$\sim 10^{-4}$
Lundquist	τ_{η} / τ_A	12	0.53	$\sim 10^4$

Measurements



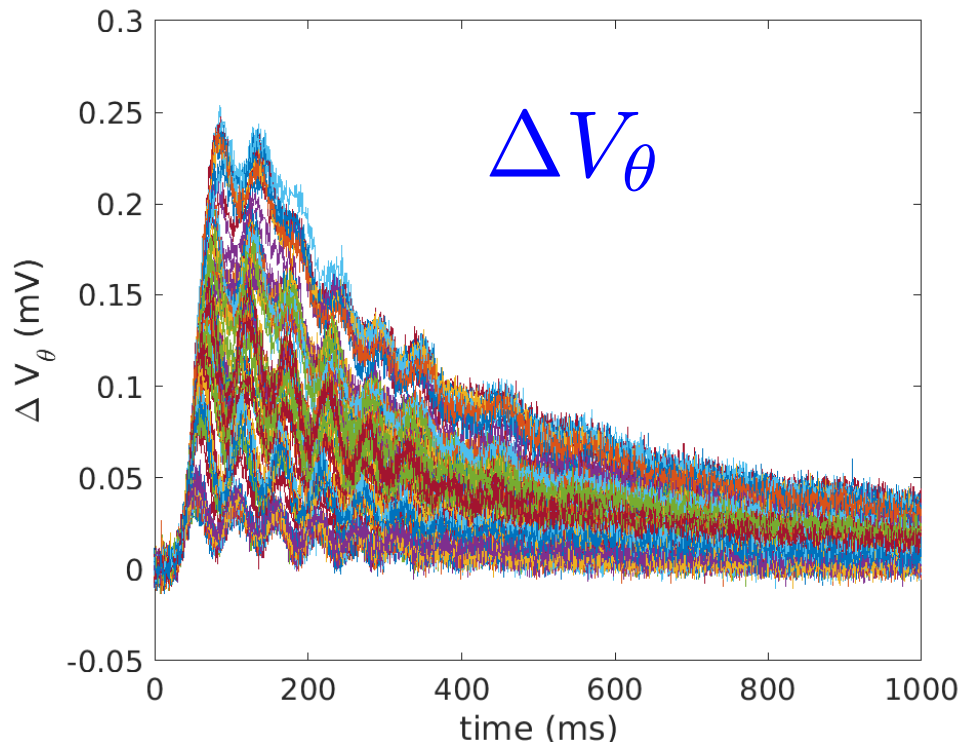
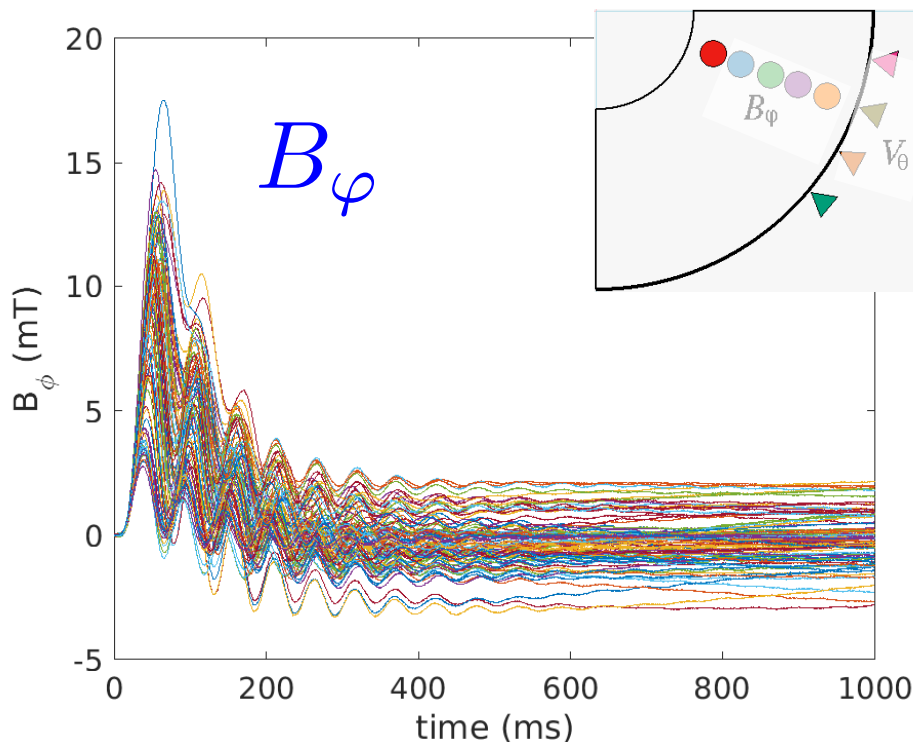
- Azimuthal magnetic field in a sleeve
- Surface electric potential
- Azimuthal fluid velocity by ultrasound Doppler

Jerks of all sizes

Azimuthal
magnetic field

$$f_o = 15 \text{ Hz}$$

Surface
electric potential





the wave comes!

*the first 80 ms
(0.16 magnetic diffusion time)*

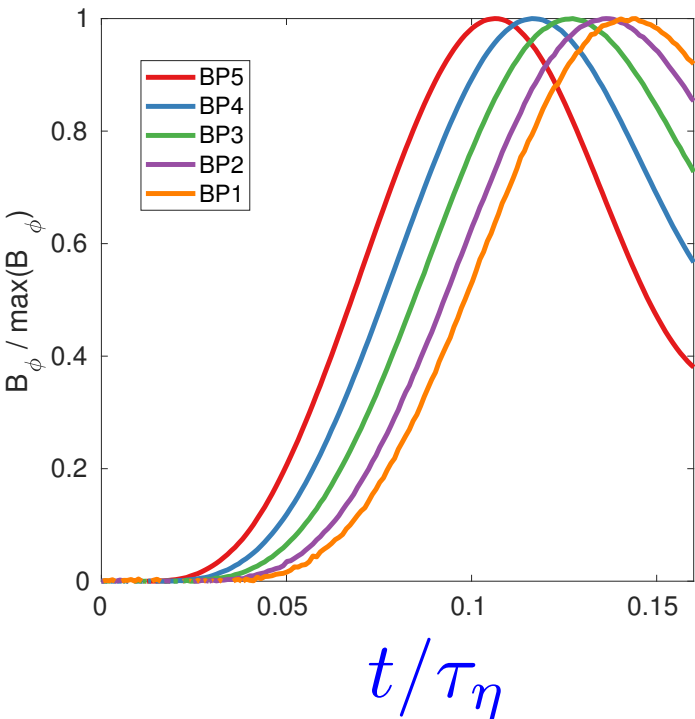
azimuthal

$$\Delta\varphi = 90^\circ, t_{rise} = 0.08\tau_\eta$$

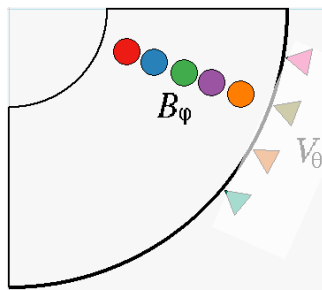
magnetic field

simulation

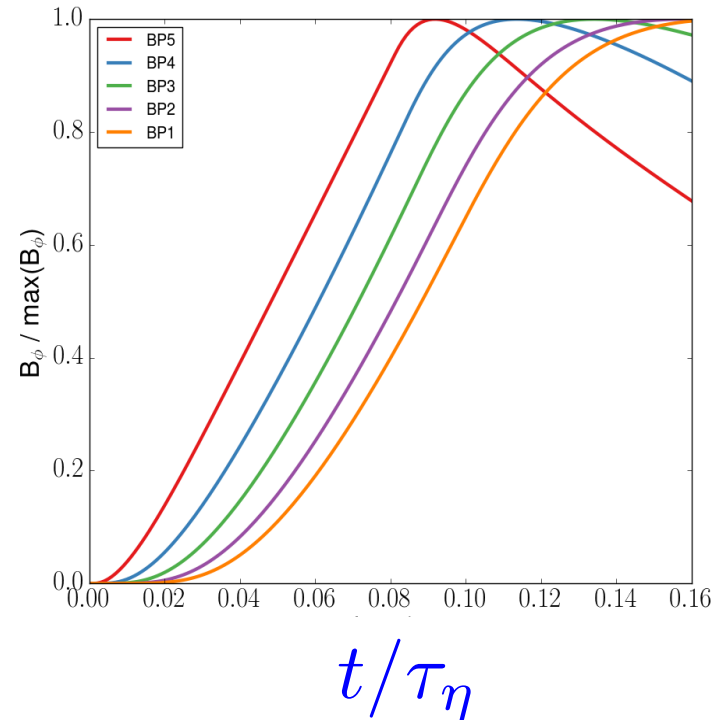
experiment



B_ϕ



$$f_o = 15 \text{ Hz}$$

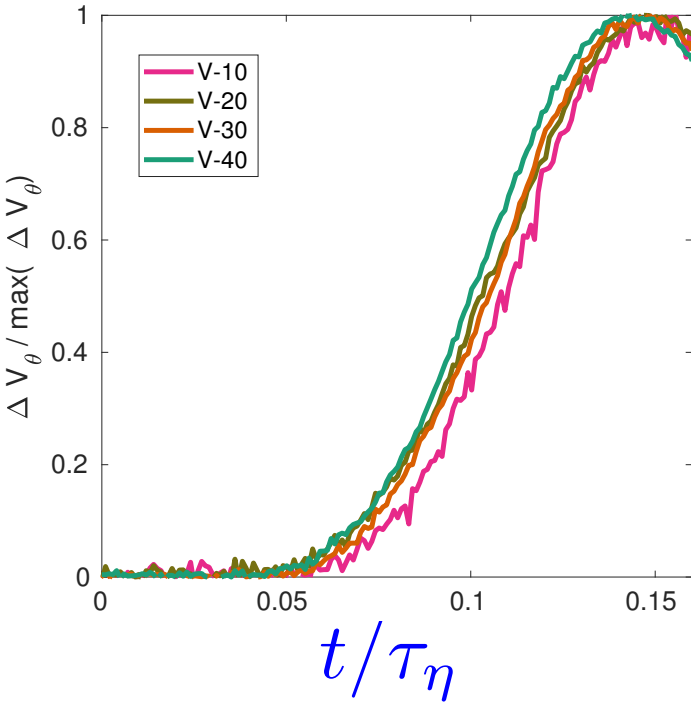


surface

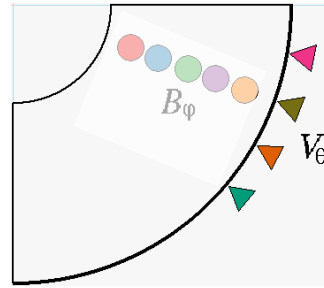
electric potential

experiment

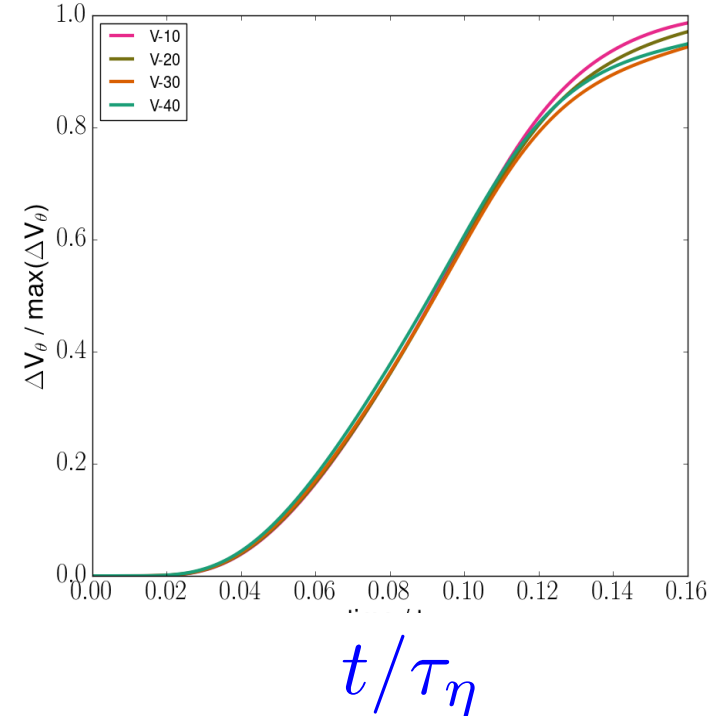
simulation



$$\Delta V_\theta$$



$$f_o = 15 \text{ Hz}$$

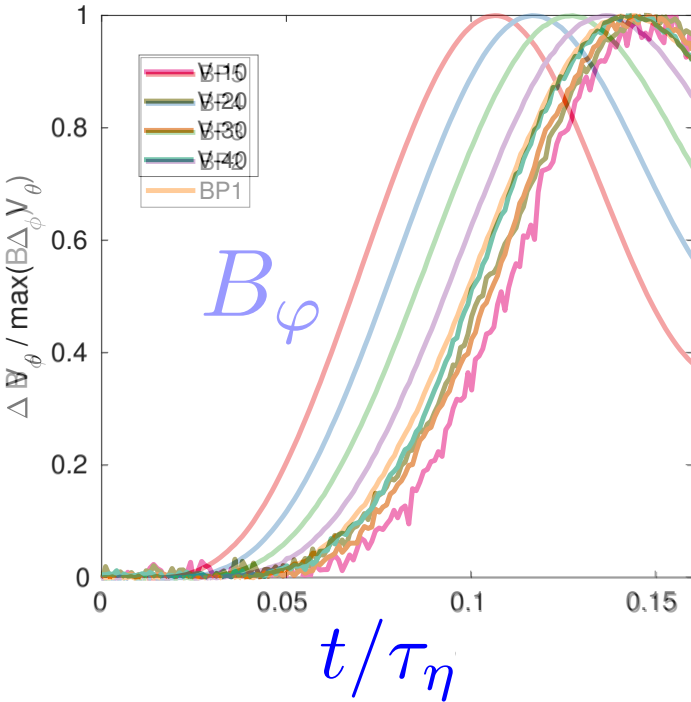


surface

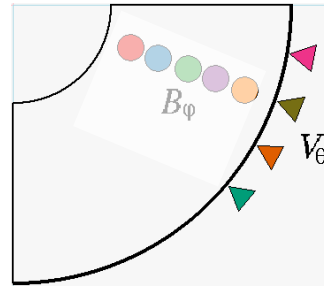
electric potential

experiment

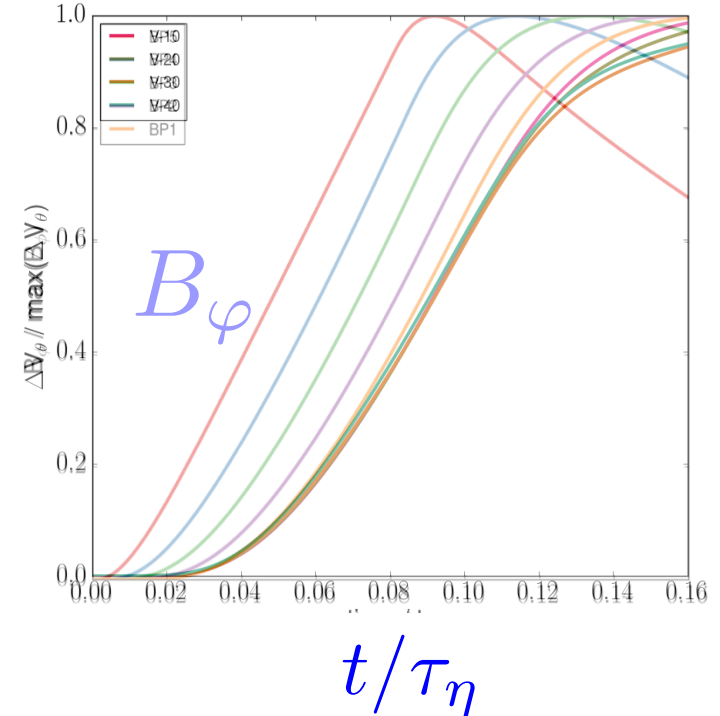
simulation



$$\Delta V_\theta$$

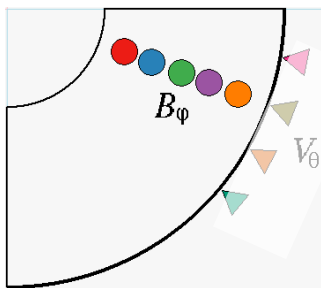


$$f_o = 15 \text{ Hz}$$





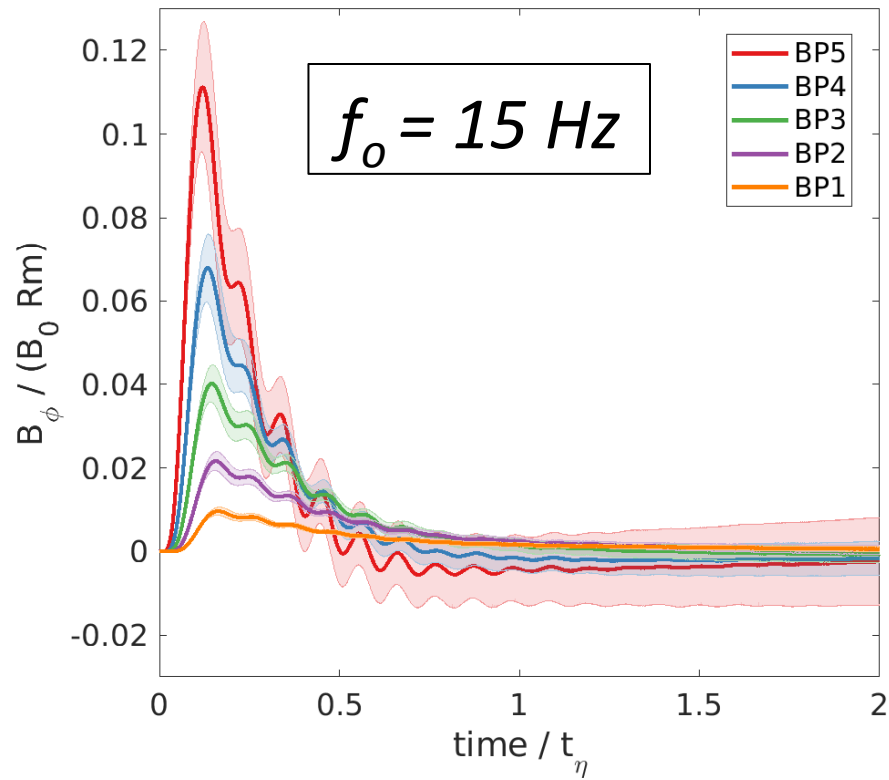
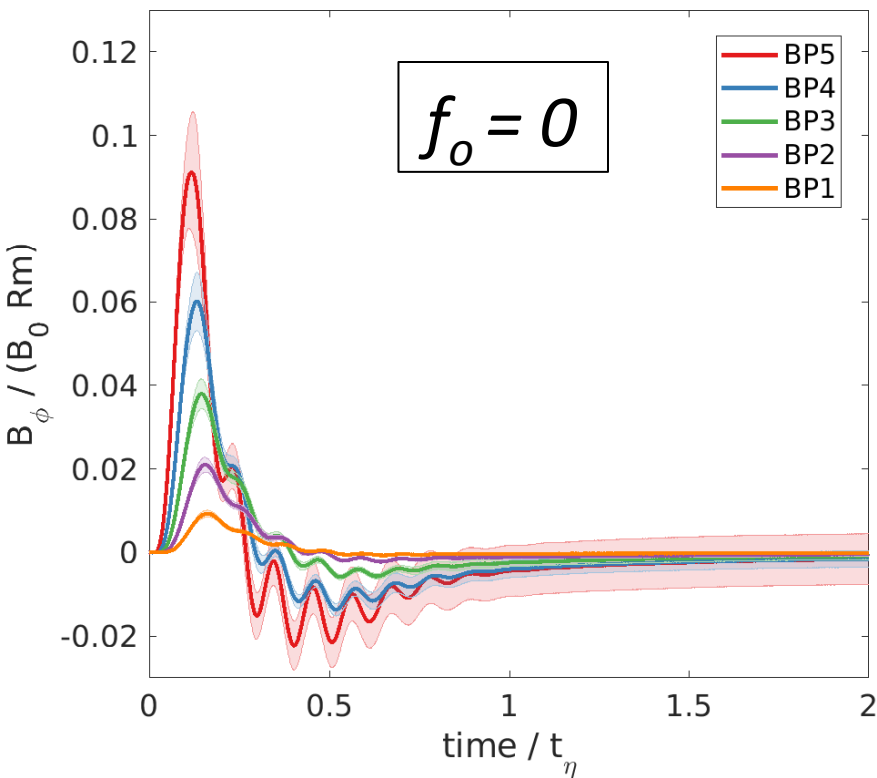
The signature of rotation:
the first 500 ms
(1 magnetic diffusion time)

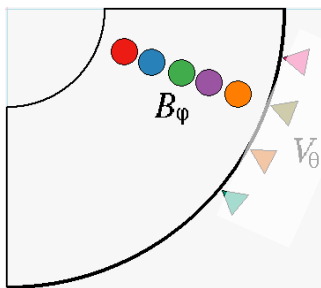


experiment

B_ϕ

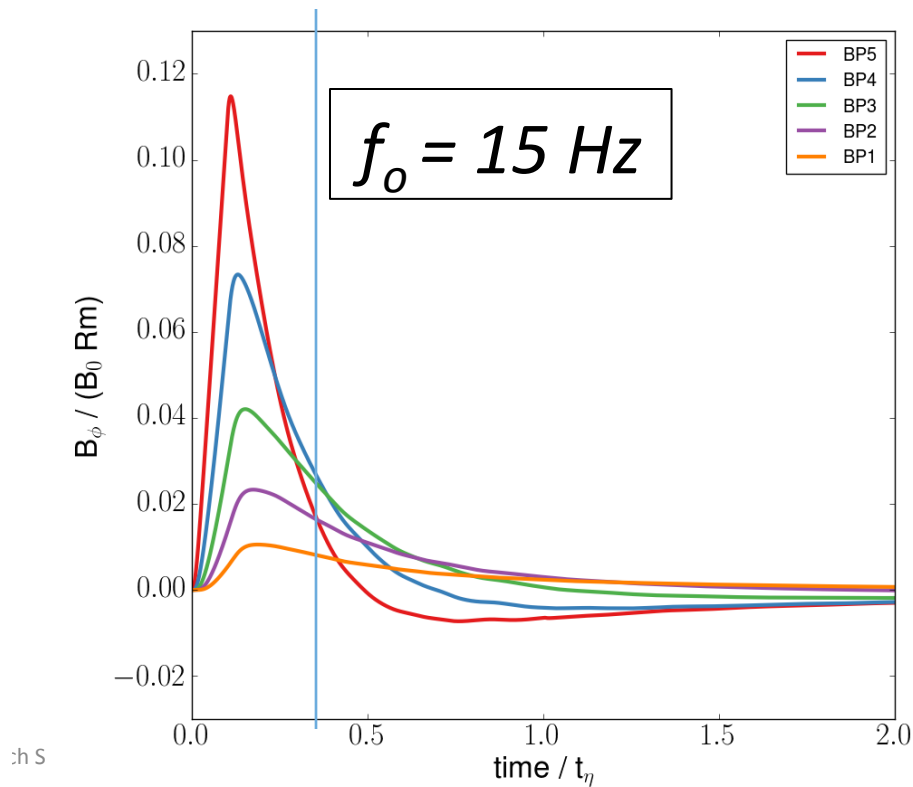
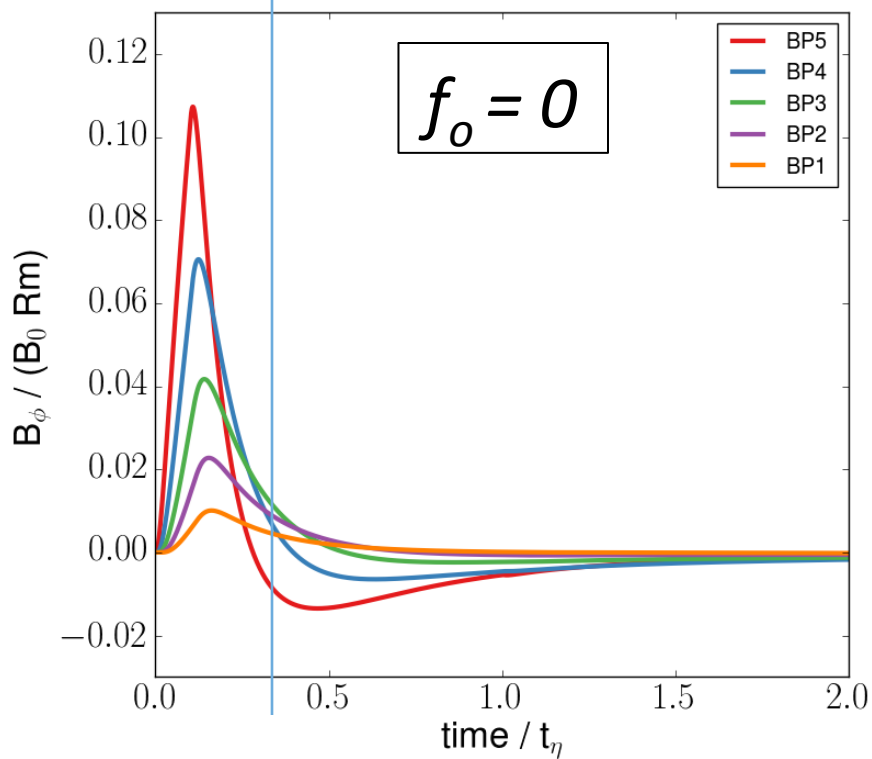
$$\Delta\varphi = 180^\circ, t_{rise} = 0.1\tau_\eta$$





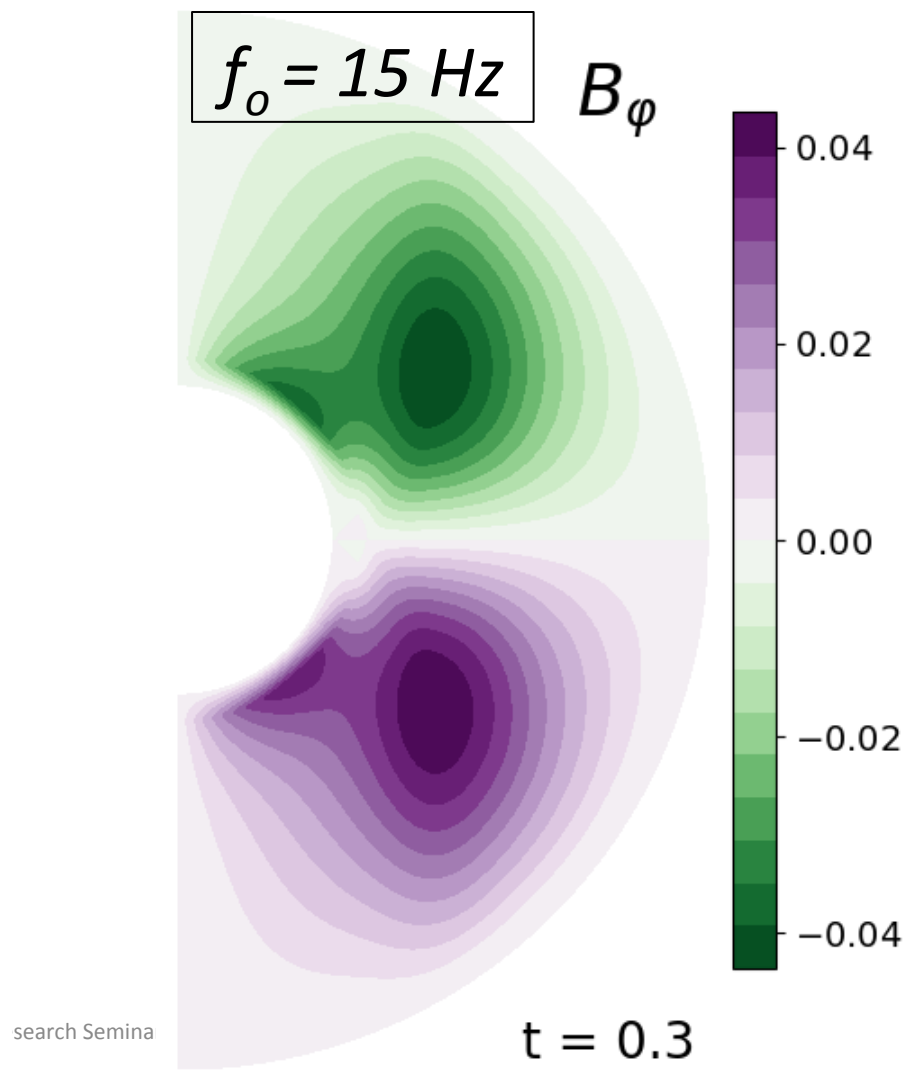
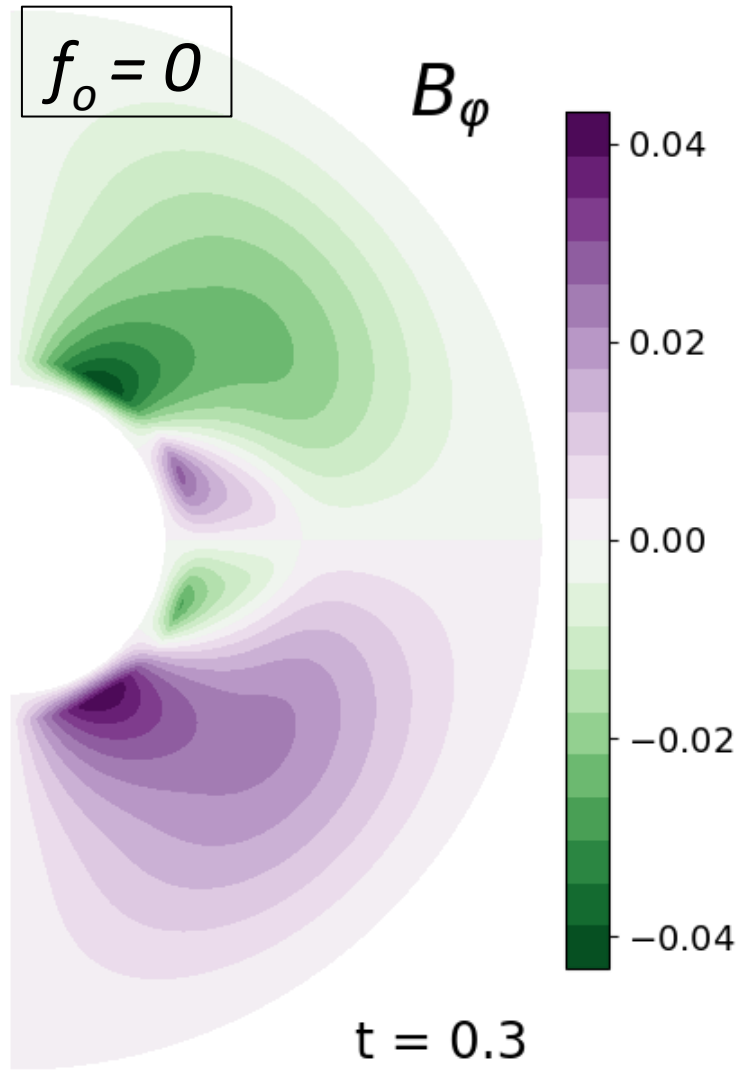
simulation

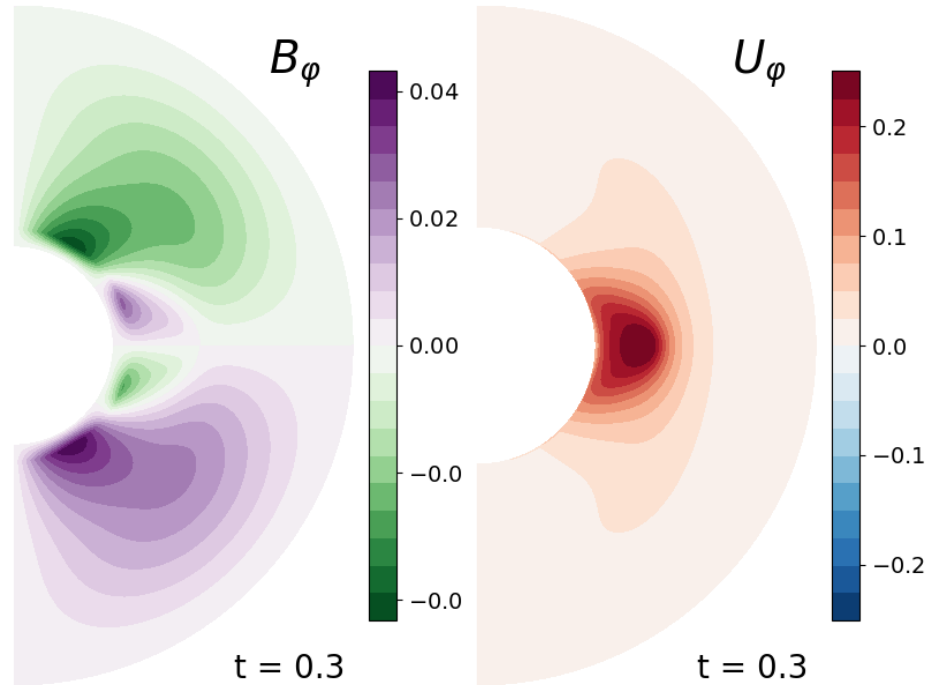
B_ϕ



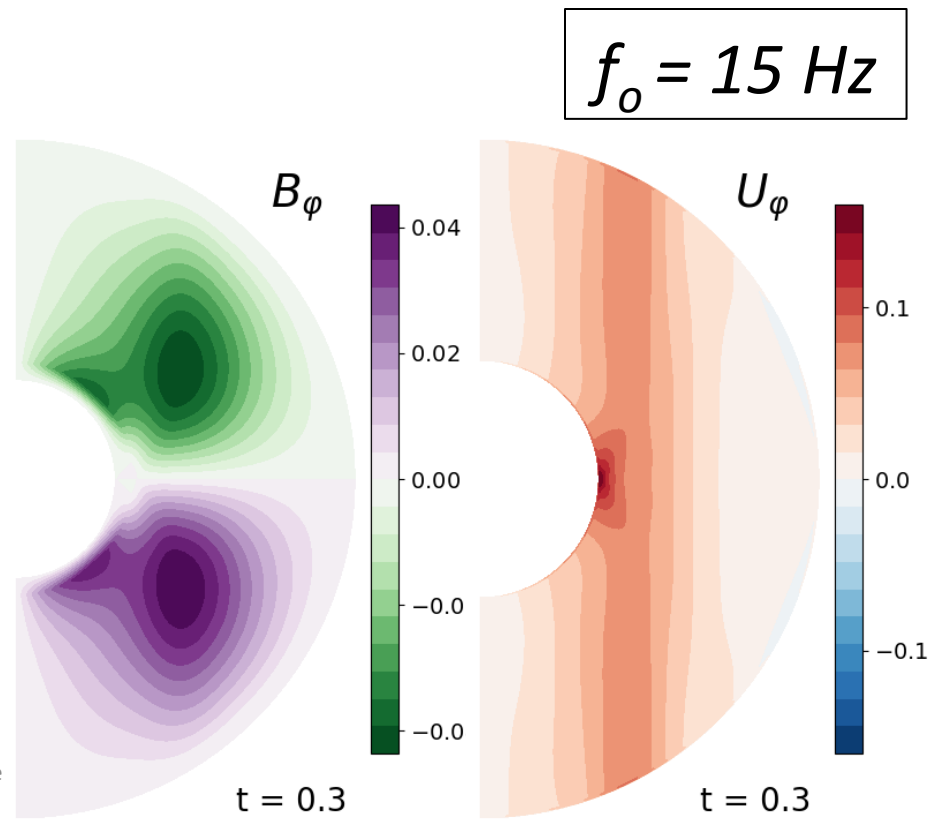


The mystery of the negative magnetic swing...



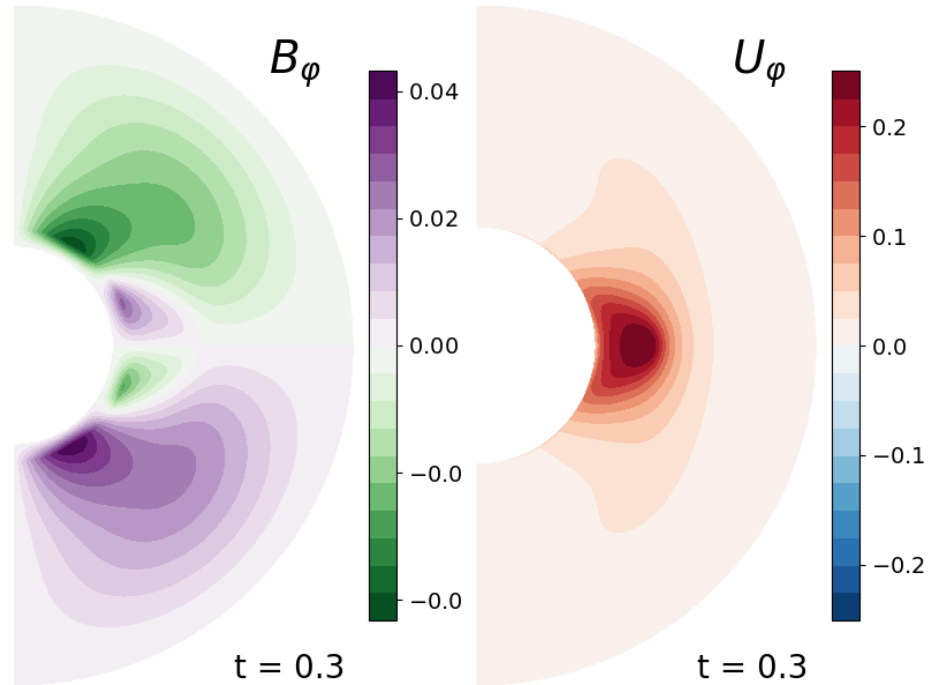


$f_o = 0$

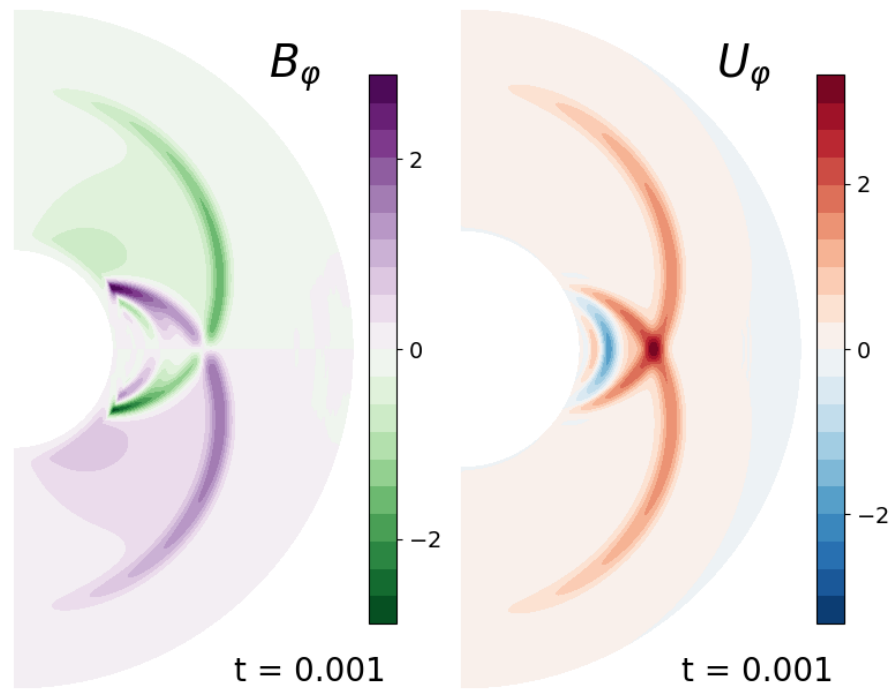


The Alfvén wave from the other side...

$$f_o = 0, Lu_i = 1200$$



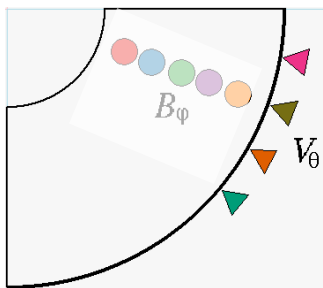
$$f_o = 0, Lu_i = 12$$





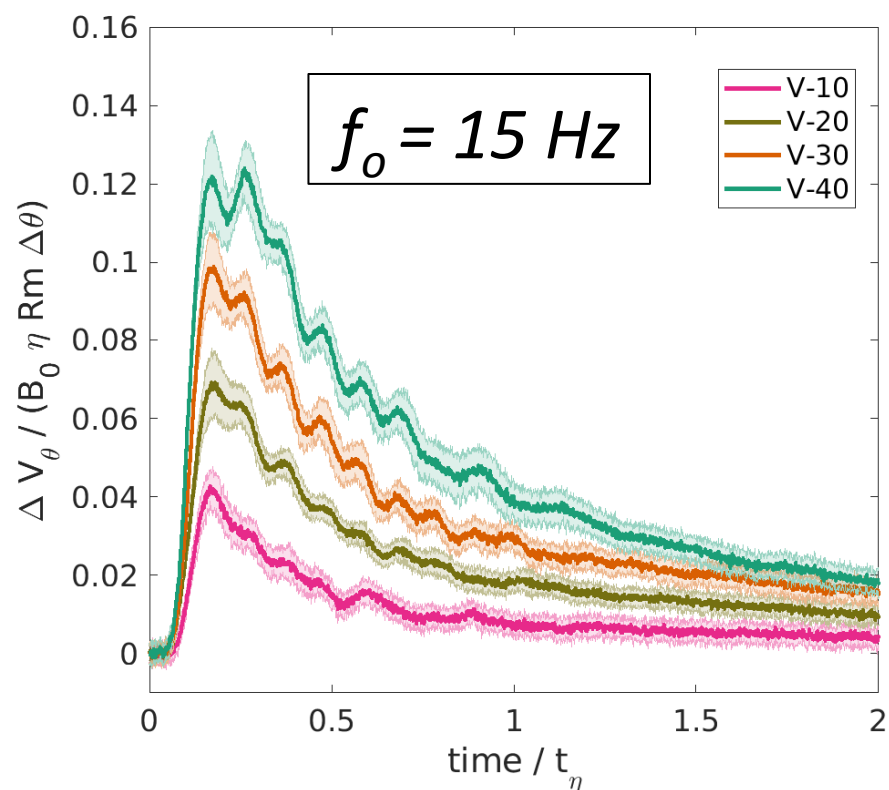
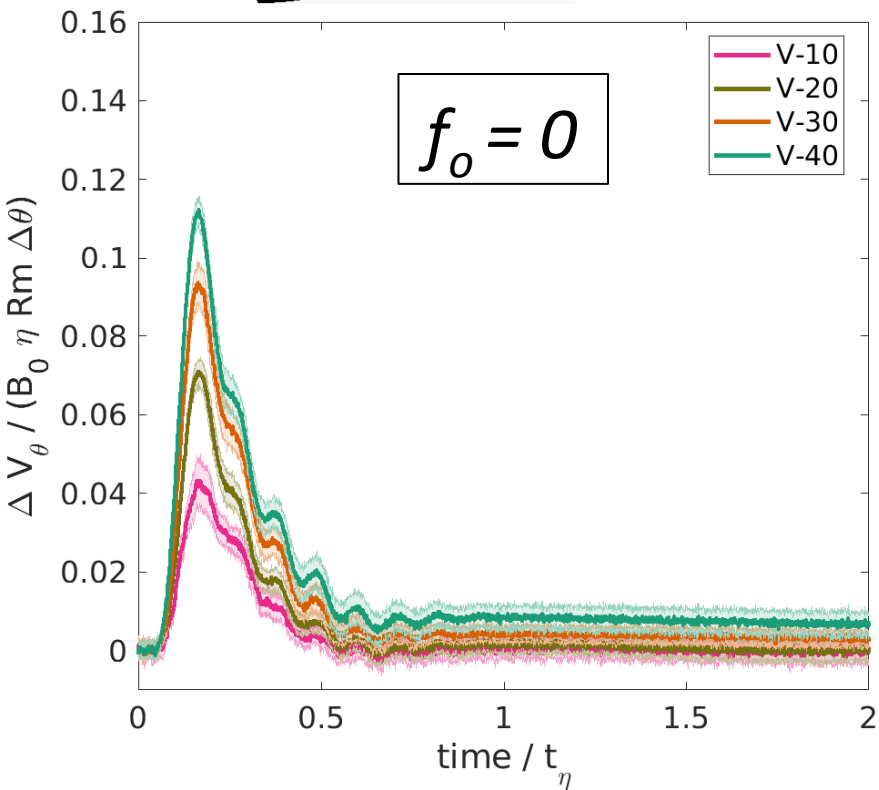
*A hint on fluid velocities from
surface electric potentials...*

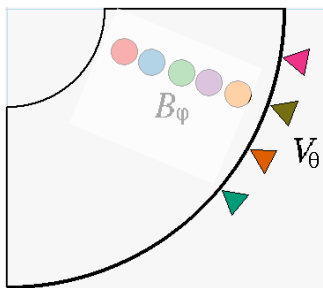
$$U_{\varphi} = \frac{1}{B_r} \frac{\Delta V_{\theta}}{r_o \Delta \theta}$$



experiment

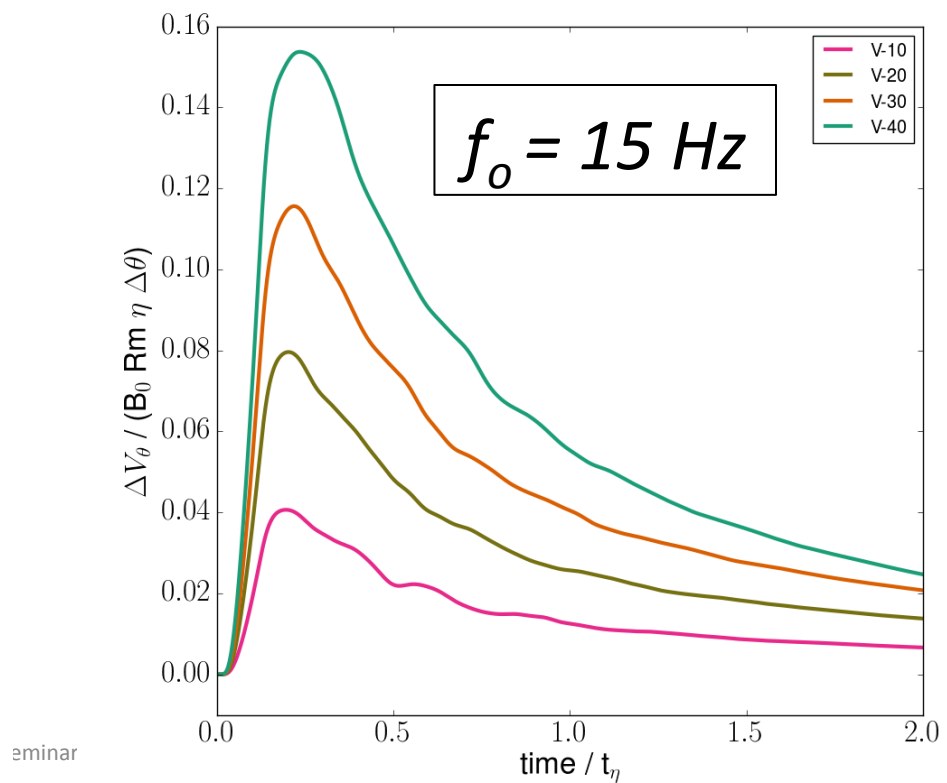
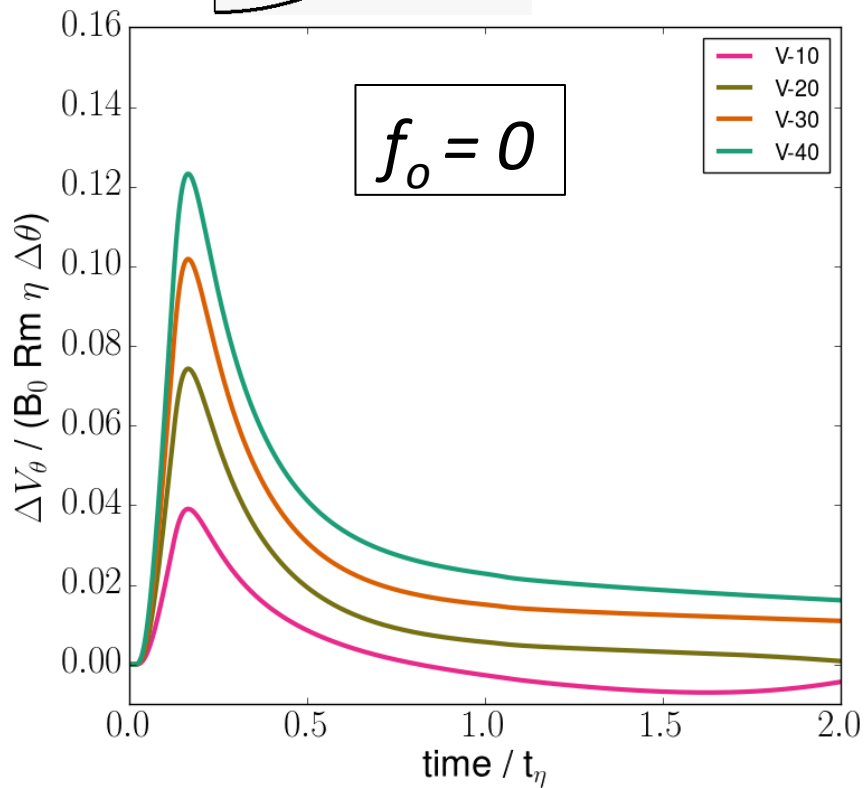
ΔV_θ

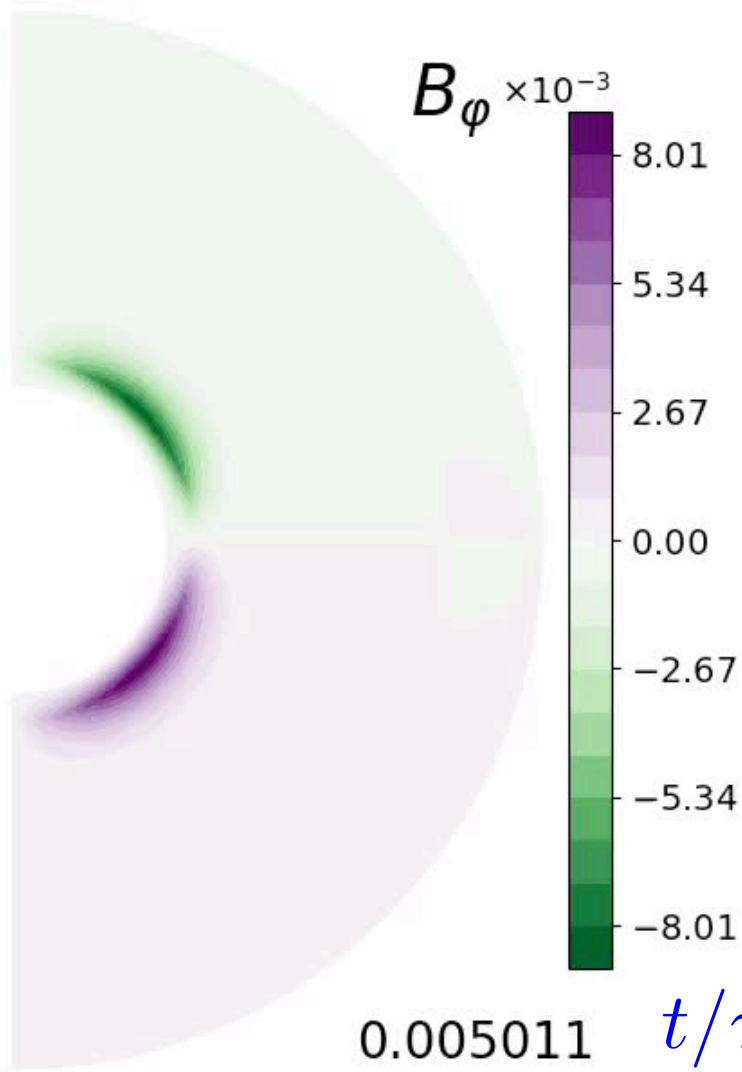




simulation

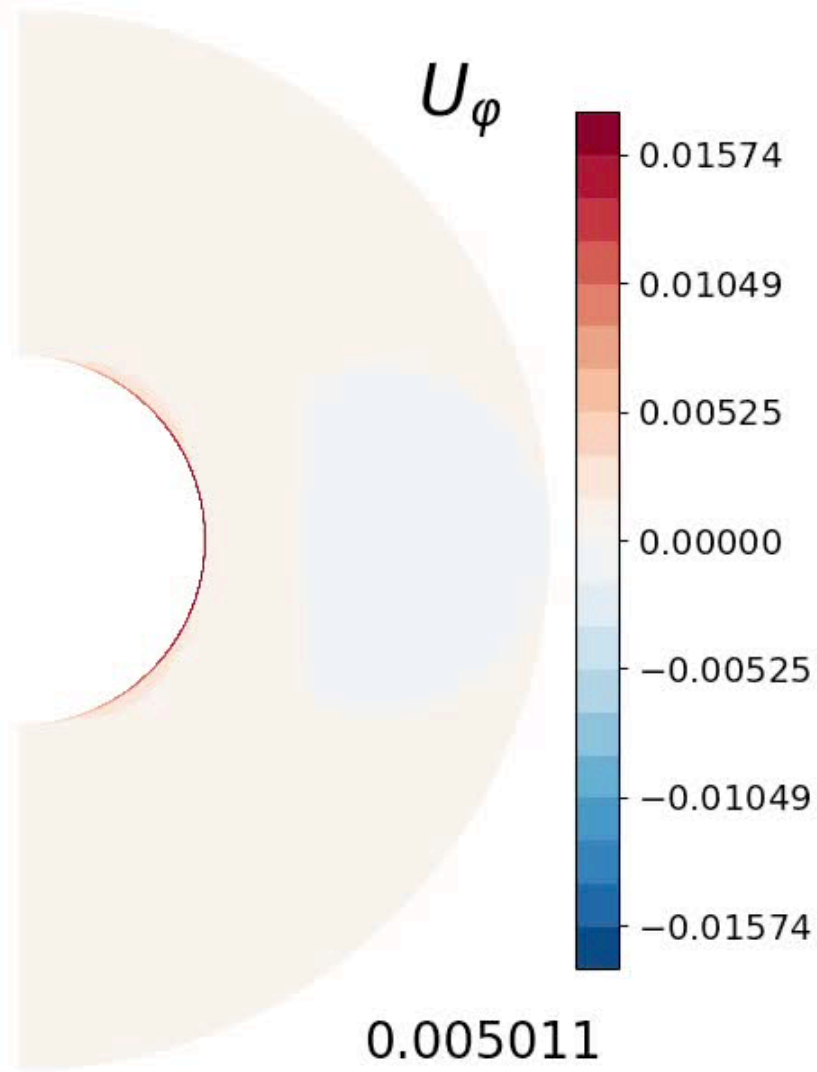
ΔV_θ





0.005011

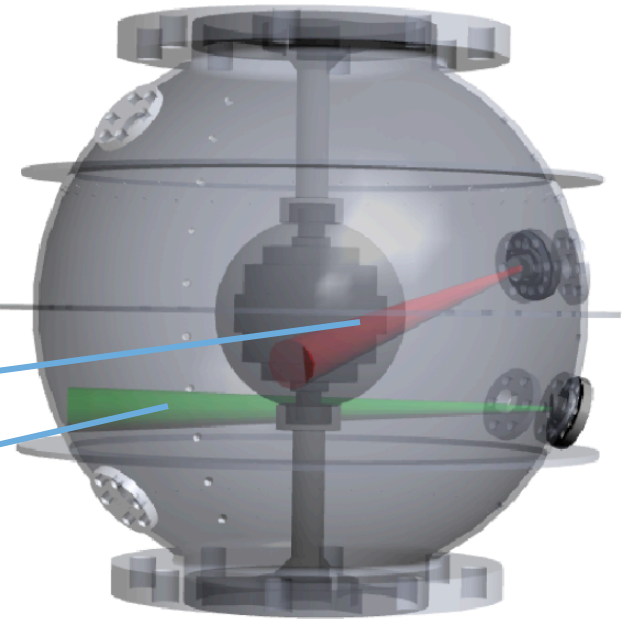
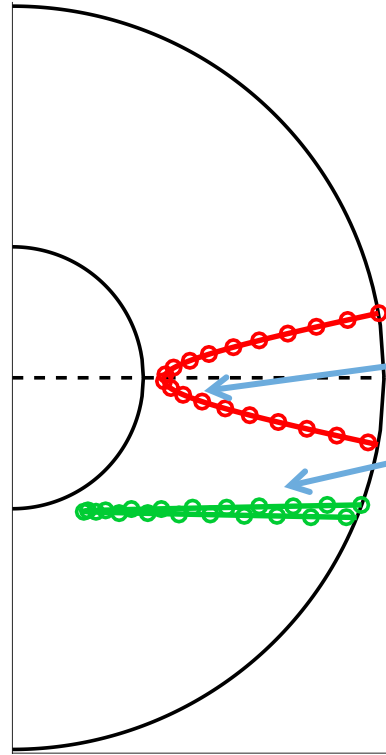
t/τ_η Research Ser



0.005011

More tricky: azimuthal velocity from UDV

Trajectory of the ultrasonic beam
in a (s,z) plane

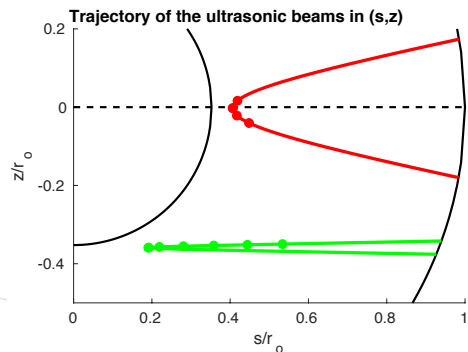
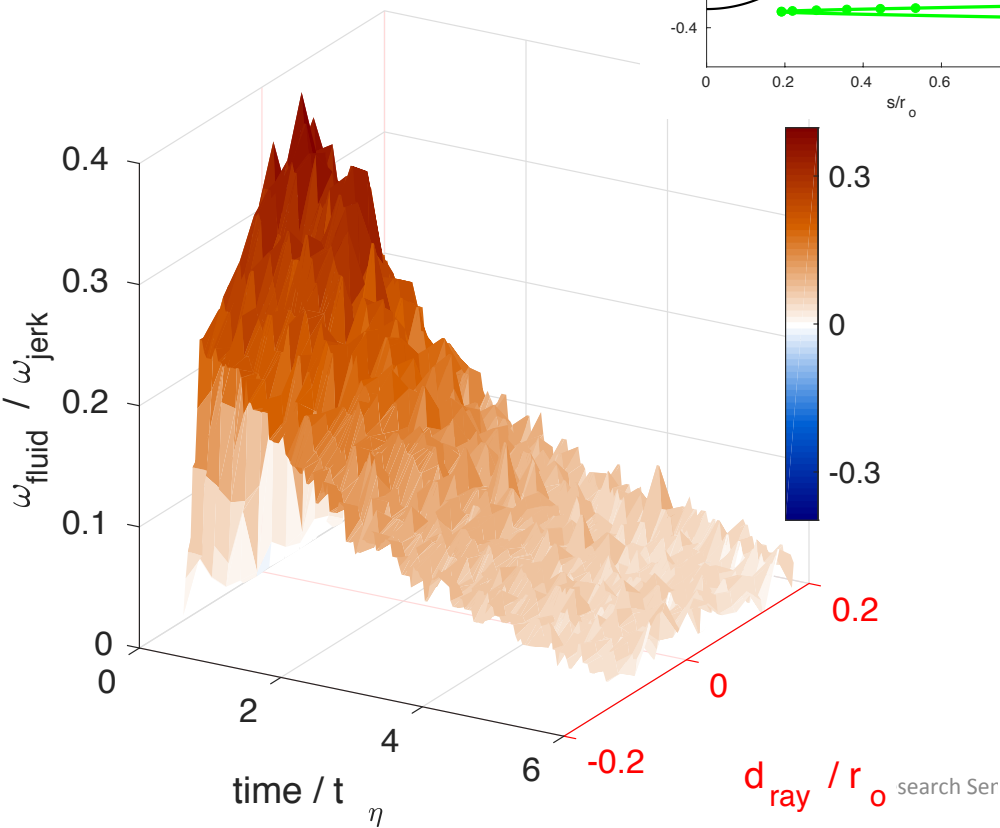


Ultrasound ray paths

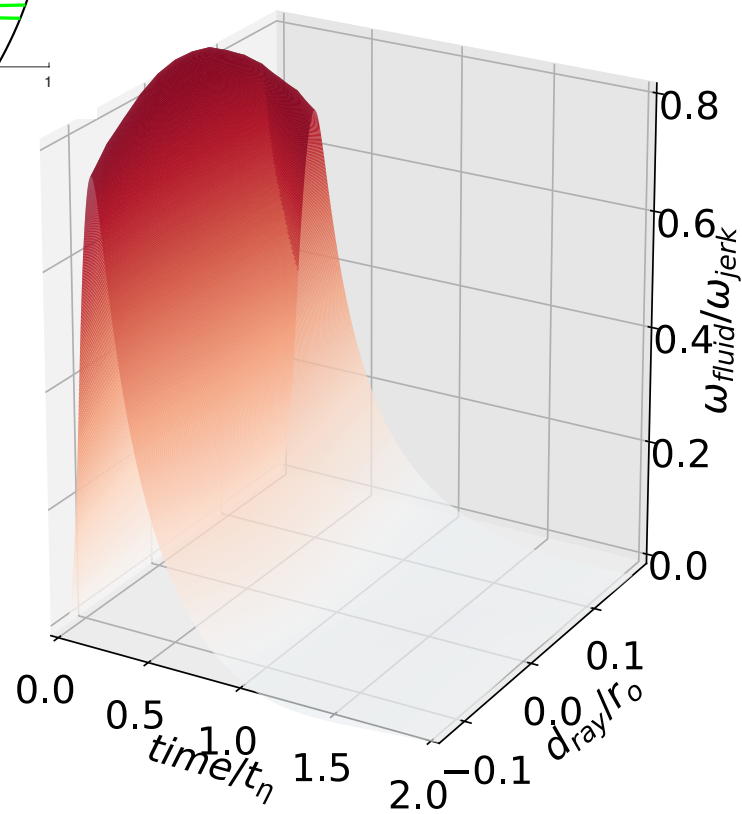


$$f_o = 0$$

experiment

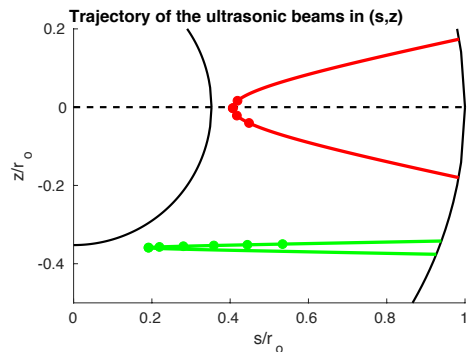
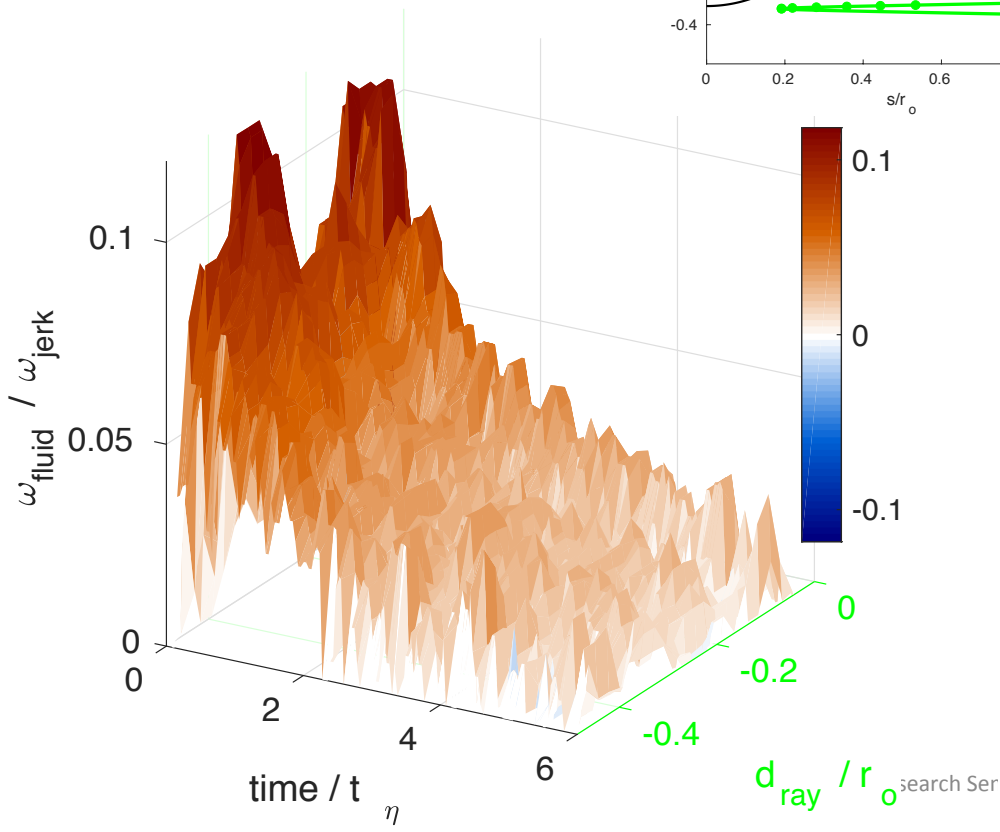


simulation

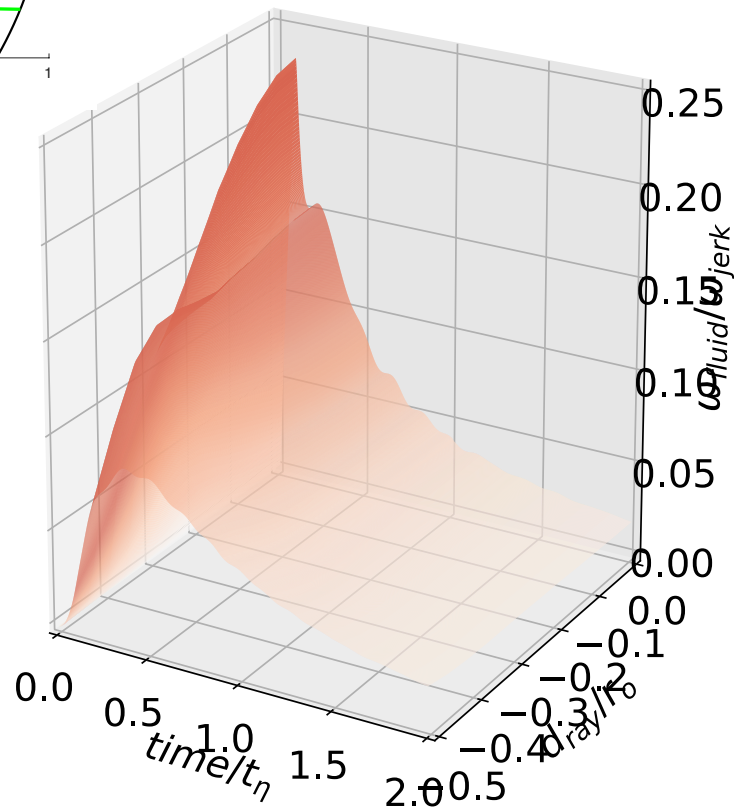


$$f_o = -10 \text{ Hz}$$

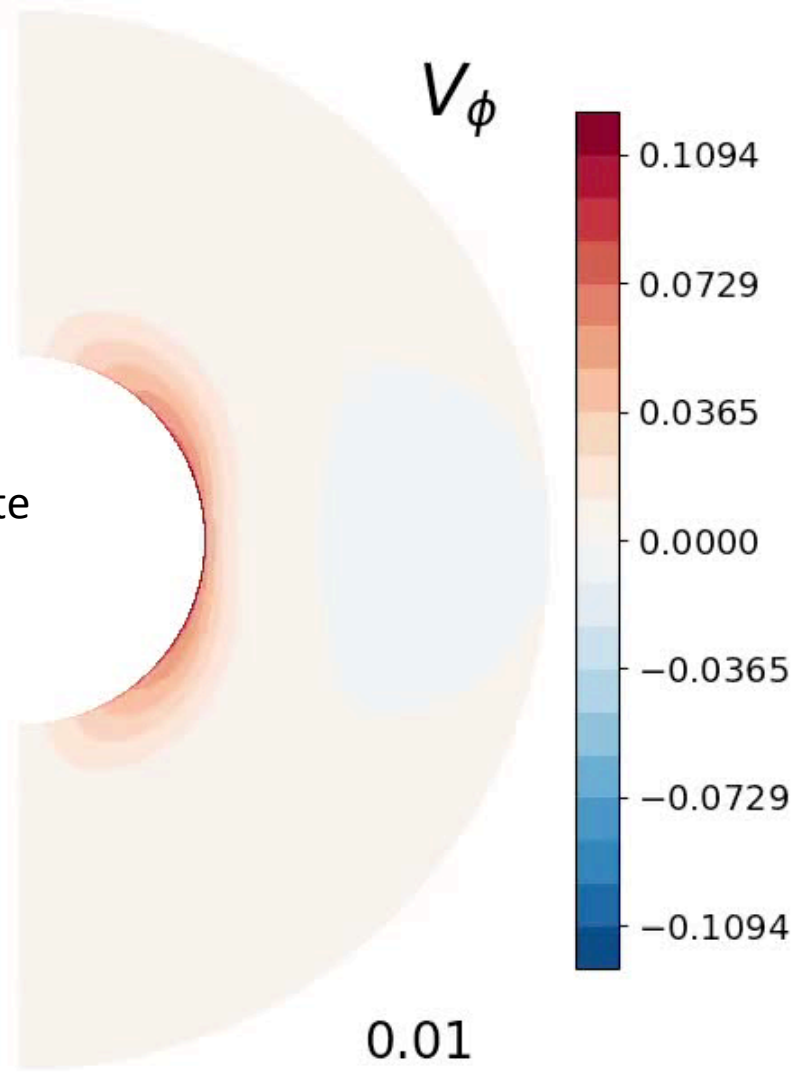
experiment



simulation

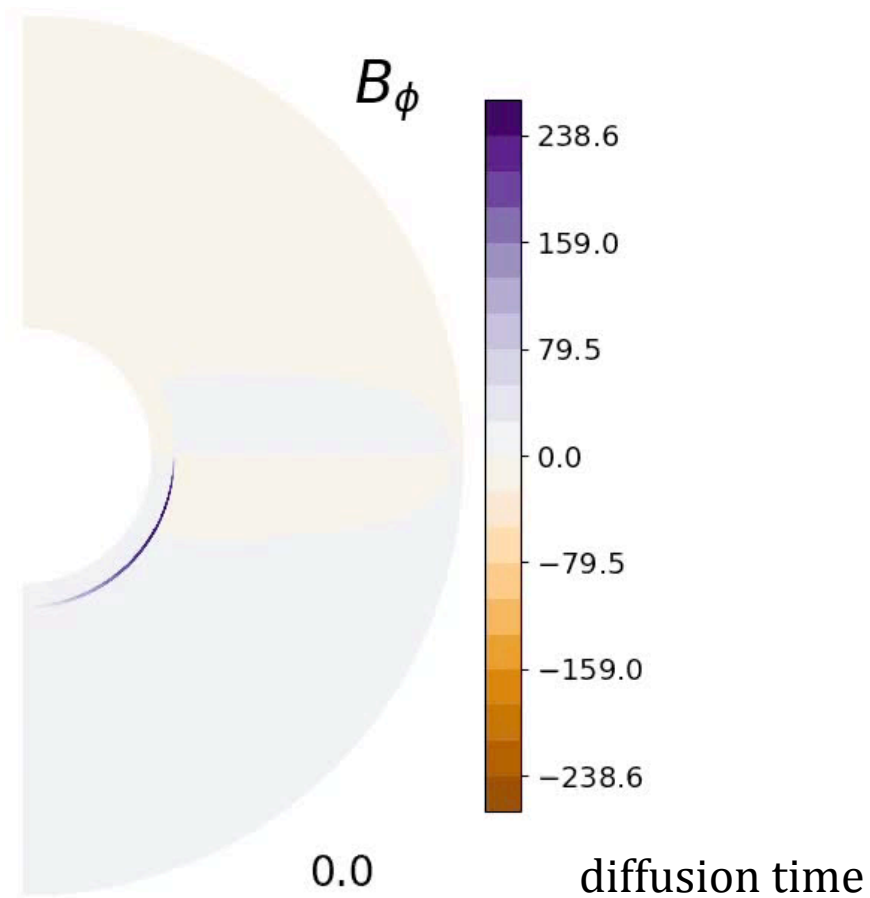


Abrupt rapid jerks excite inertial modes

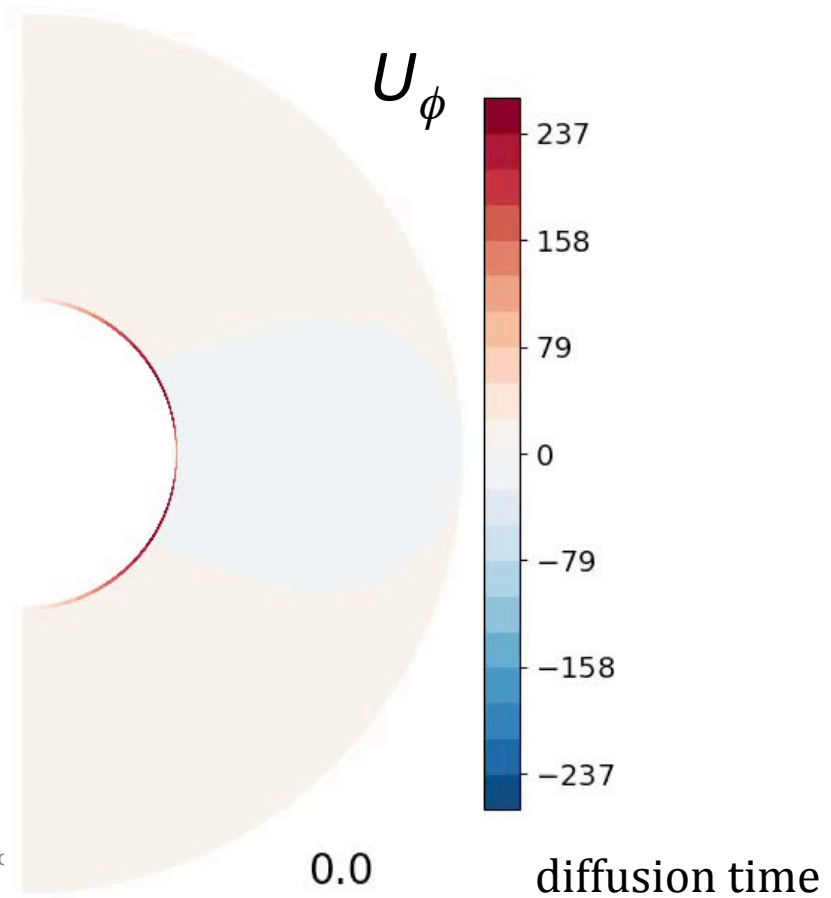


More Earth-like values:

$\text{Lu}_i=12\,000$, $\text{Le}_o=10^{-3}$
 $E=2\times 10^{-7}$, $\text{Ro}=2\times 10^{-2}$ (but $\text{Pm}=0.1$)



search Seminar, Kyc

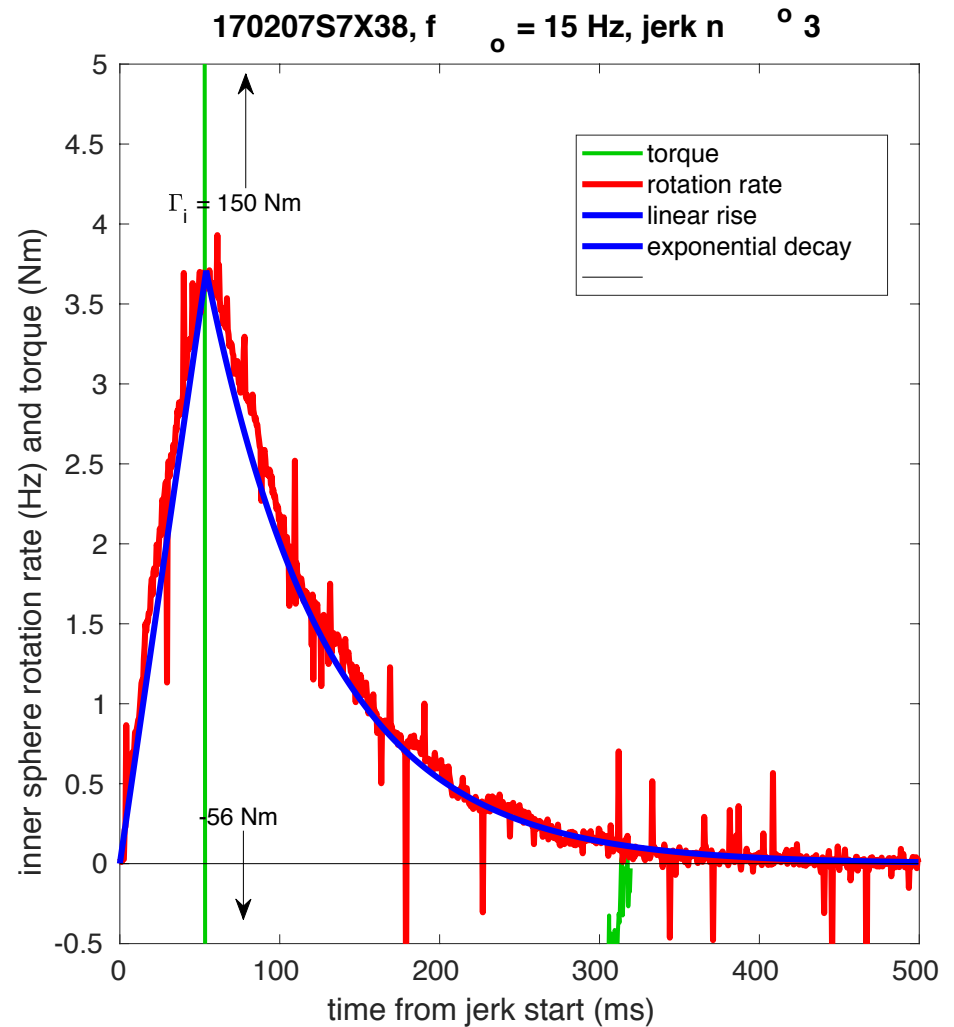


Take-home message

- We have triggered and observed **torsional Alfvén waves** in our DTS Ω laboratory experiment.
- Rotation, magnetic field geometry and diffusion **strongly alter** ideal Alfvén wave properties.
- XSHELLS numerical simulations help deciphering their properties, and show the triggering of **inertial waves**.
- Electric potentials and subtle differences in the magnetic signature reveal the formation of **geostrophic motions**.
- We obtained the first direct measurements of **Alfvén wave fluid velocity** from ultrasound Doppler.

Thank you

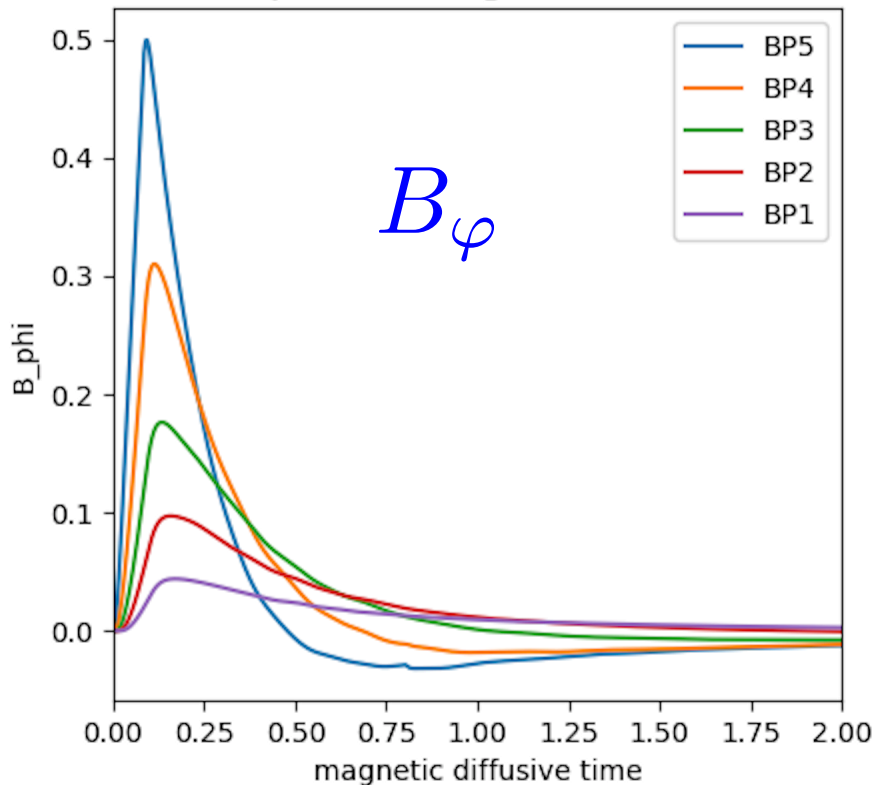
DTSΩ jerk time function



simulation

$$f_0 = 15 \text{ Hz}$$

Bphi at -20 deg versus time



Uphi at -20 deg versus time

